

.. NLO: How to?

# IWATE COLLIDER SCHOOL 2024

**26 FEBRUARY - 2 MARCH, 2024**

Appi highland, Iwate, Japan



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# Introduction:

## Why do we need $N^{(k)}$ LO?

*why?*

**why?**

**why?**

*why?*

*why?*

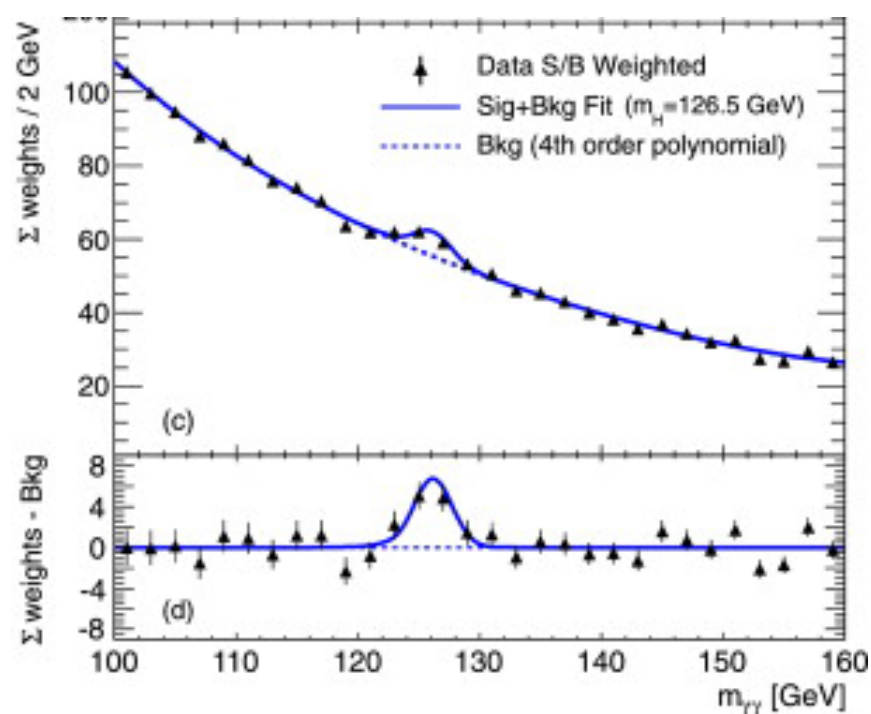


# Discoveries at hadron colliders

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## Peak

$$H \rightarrow \gamma\gamma$$



**EASY**

Background directly measured  
from **data**.

Theory needed only for  
parameter extraction



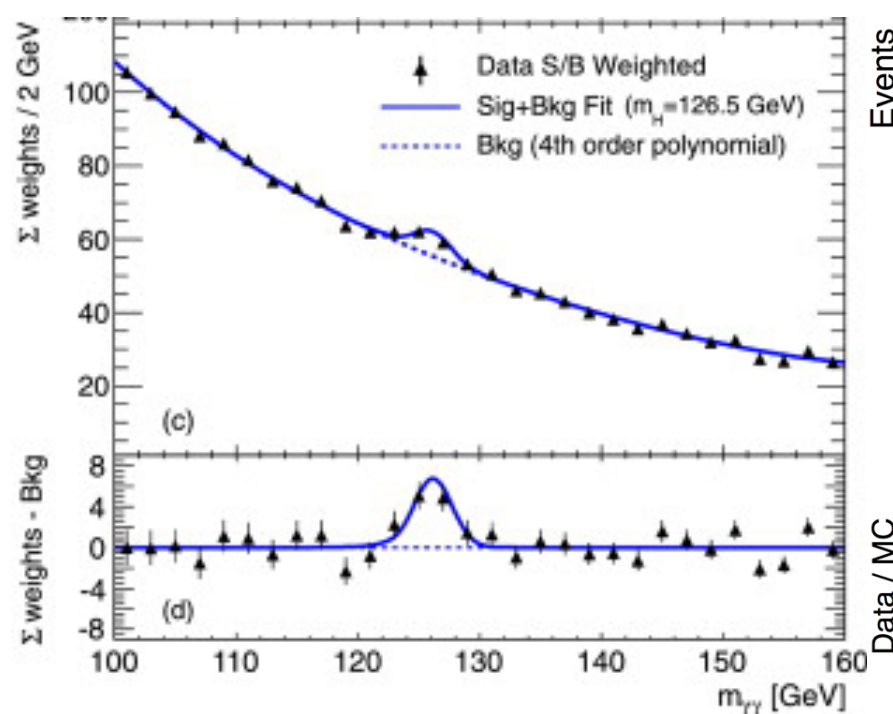
# Discoveries at hadron colliders

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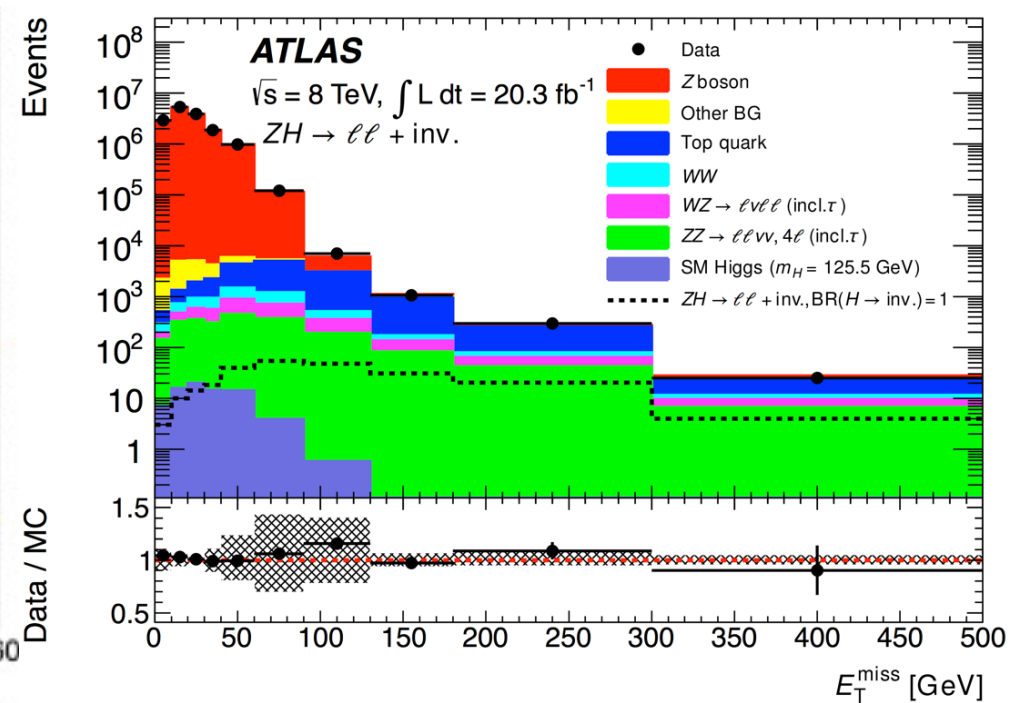
$$ZH \rightarrow l^+l^- + inv.$$



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Background **SHAPE** needed. Flexible MC for both signal and background validated and tuned to data

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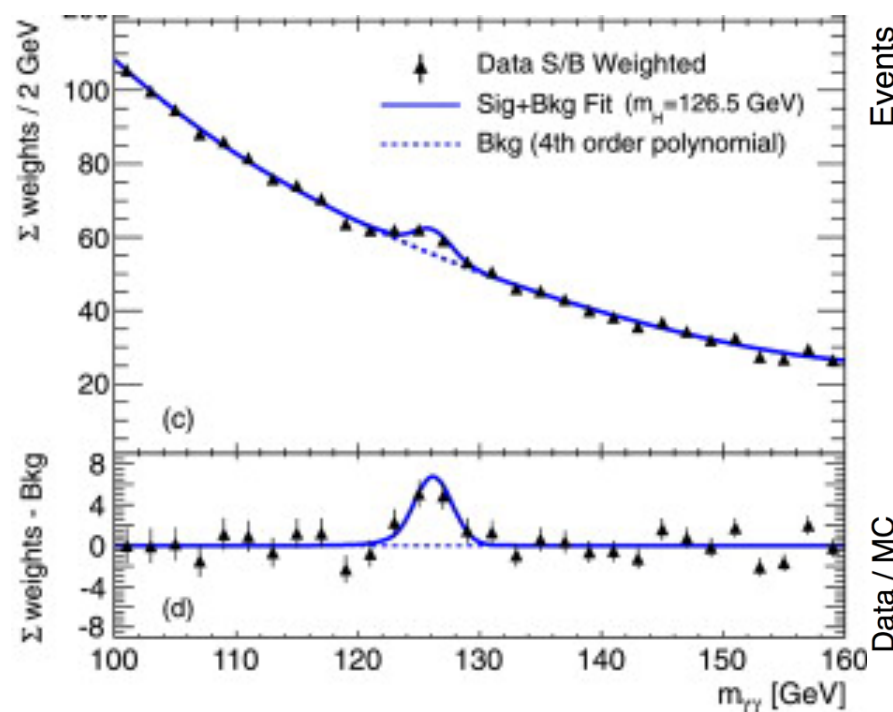
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## Rate

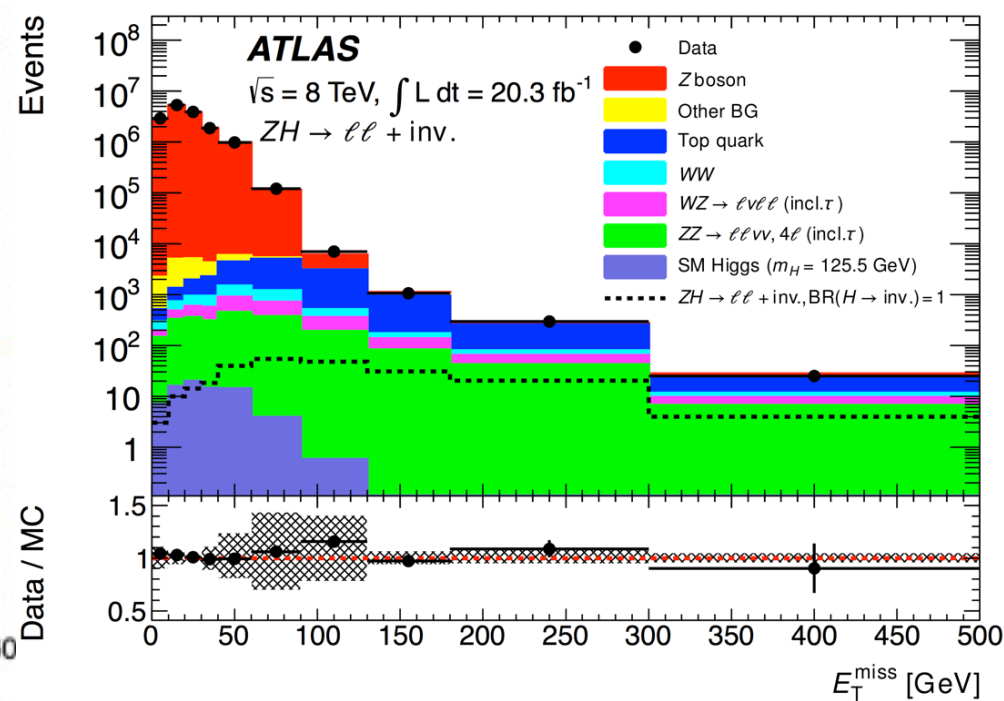
$$H \rightarrow W^+ W^-$$



## EASY

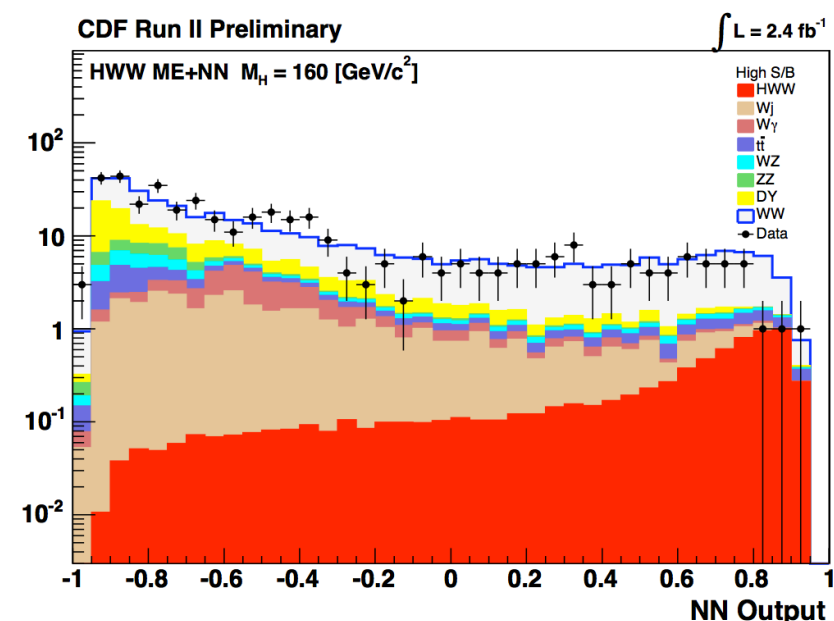
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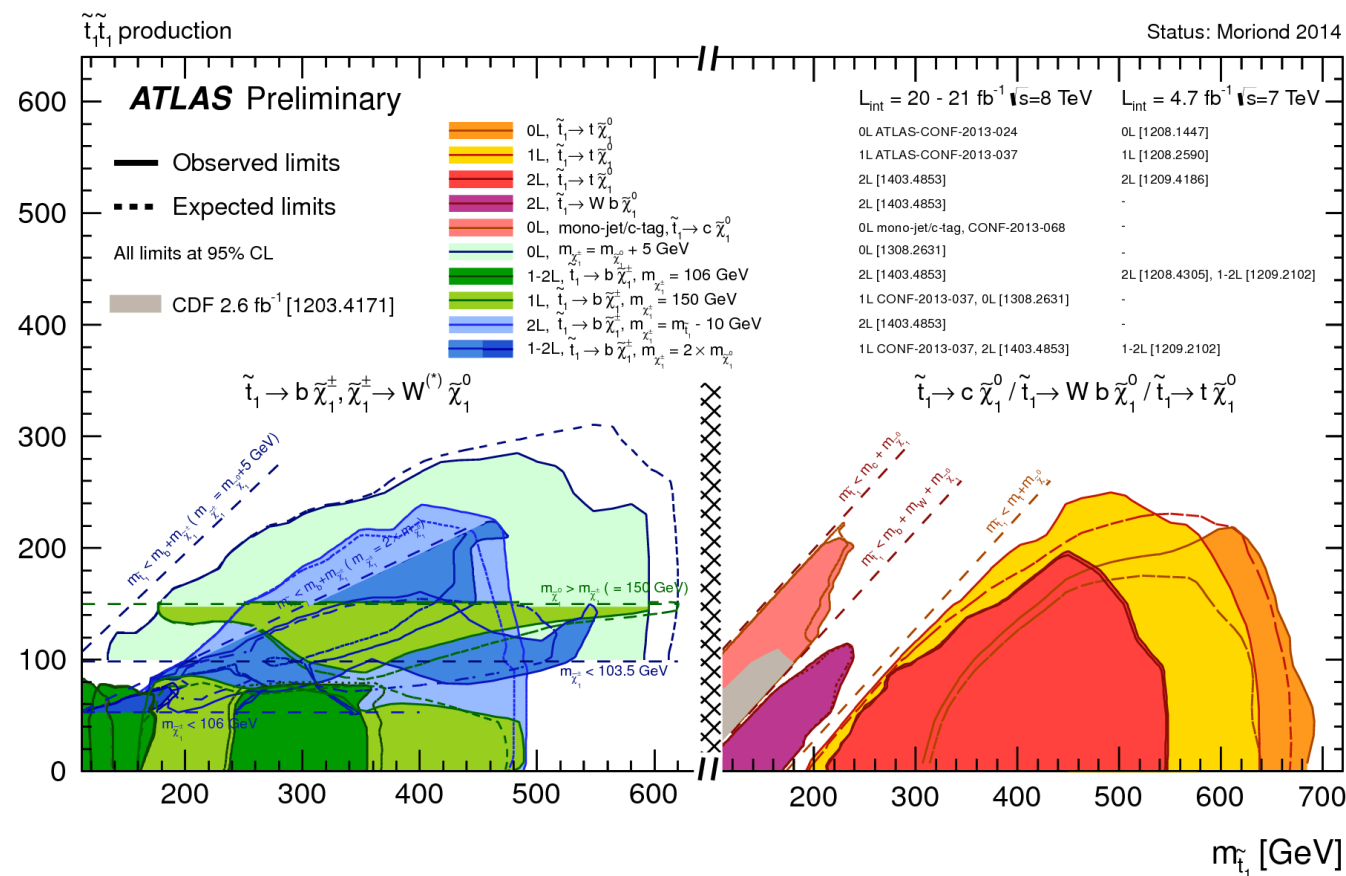


## VERY HARD

Relies on prediction for both **shape** and **normalization**. Complicated interplay of best simulations and data

# New physics?

- No NP has been discovered yet
- Either there is no NP, or it is hiding very well
- If it is there, it will be a 'Hard' or 'very Hard' discovery
- Need for accurate predictions for signal and background



ATLAS SUSY Searches\* - 95% CL Lower Limits  
Status: Moriond 2014

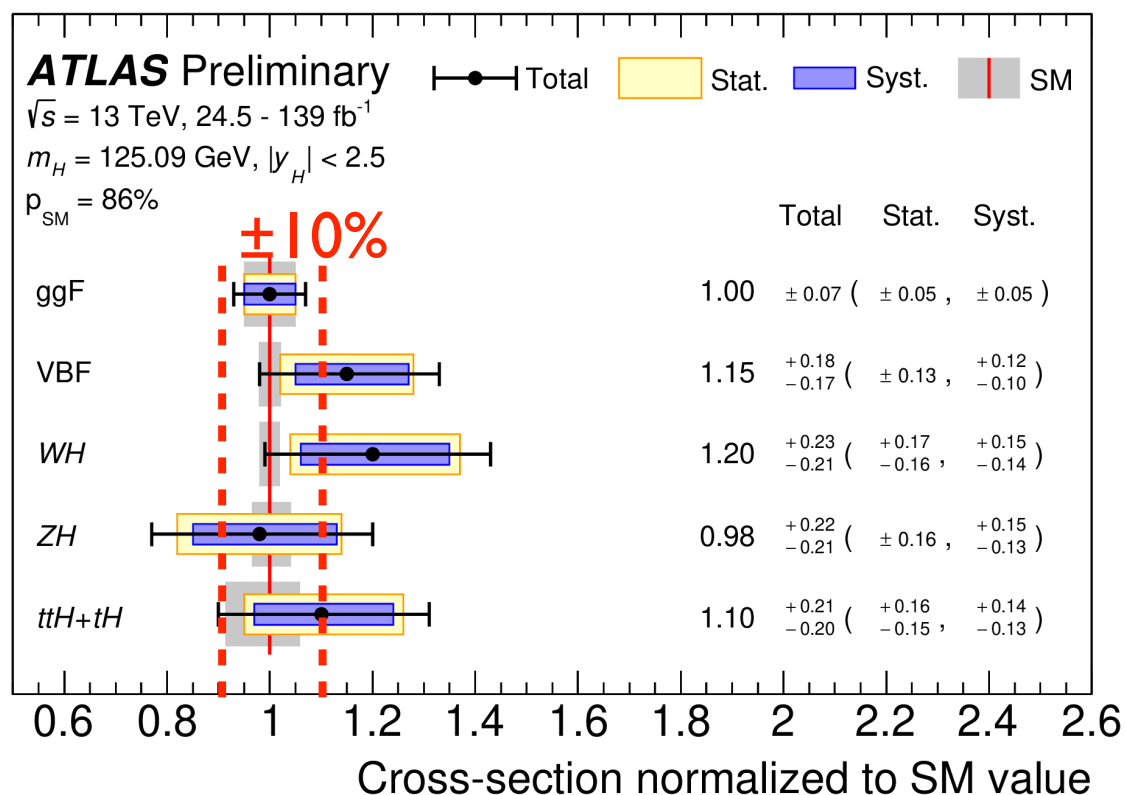
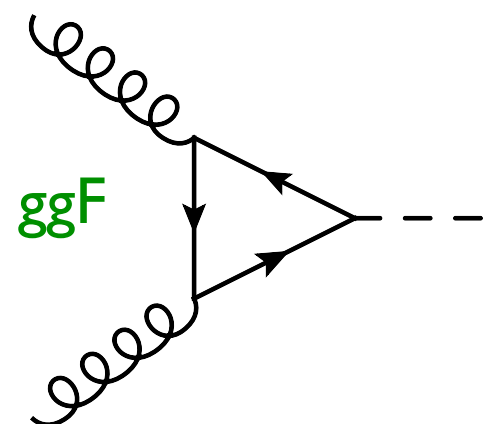
ATLAS Preliminary  
 $\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1}$   $\sqrt{s} = 7, 8 \text{ TeV}$

Model	$e, \mu, \tau, \gamma$	Jets	$E_{\text{miss}}$	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Mass limit	Reference		
Inclusive Searches	MSUGRA/CMSSM	0	2-6 jets	Yes	20.3	$\tilde{g}, \tilde{t}_1$ 1.7 TeV	ATLAS-CONF-2013-047	
	MSUGRA/CMSSM	$1 e, \mu$	3-6 jets	Yes	20.3	any $m(\tilde{g})$	ATLAS-CONF-2013-062	
	MSUGRA/CMSSM	0	7-10 jets	Yes	20.3	any $m(\tilde{g})$	1308.1841	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}$	0	2-6 jets	Yes	20.3	$m(\tilde{g}) = 0 \text{ GeV}$	ATLAS-CONF-2013-047	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q} \ell_1$	0	2-6 jets	Yes	20.3	$m(\tilde{g}) = 0 \text{ GeV}$	ATLAS-CONF-2013-047	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q} \ell_1 \rightarrow q\bar{q} W \nu \tilde{\chi}_1^0$	$1 e, \mu$	3-6 jets	Yes	20.3	$m(\tilde{g}) < 200 \text{ GeV}, m(\tilde{g}) = 0.5(m(\tilde{g}) + m(\tilde{g}))$	ATLAS-CONF-2013-062	
	GMSB ( $\tilde{g}$ NLSP)	$2 e, \mu$	0-3 jets	-	20.3	$m(\tilde{g}) = 0 \text{ GeV}$	ATLAS-CONF-2013-089	
	GMSB ( $\tilde{g}$ NLSP)	$2 e, \mu$	2-4 jets	Yes	4.7	$\tan\beta < 15$	1208.4688	
	GGM (bino NLSP)	$1-2 \tau$	0-2 jets	Yes	20.7	$\tan\beta > 18$	ATLAS-CONF-2013-026	
	GGM (wino NLSP)	$2 \gamma$	-	Yes	20.3	$m(\tilde{g}) > 50 \text{ GeV}$	ATLAS-CONF-2014-001	
$3^{\text{rd}}$ gen. $\tilde{g}$ med.	$\tilde{g} \rightarrow b\bar{b}$	0	3 b	Yes	20.1	$m(\tilde{g}) > 50 \text{ GeV}$	ATLAS-CONF-2013-061	
	$\tilde{g} \rightarrow t\bar{t}$	0	7-10 jets	Yes	20.3	$m(\tilde{g}) < 350 \text{ GeV}$	1308.1841	
	$\tilde{g} \rightarrow c\bar{c}$	0-1 $e, \mu$	3 b	Yes	20.1	$m(\tilde{g}) < 400 \text{ GeV}$	ATLAS-CONF-2013-061	
	$\tilde{g} \rightarrow b\bar{b}$	0-1 $e, \mu$	3 b	Yes	20.1	$m(\tilde{g}) < 300 \text{ GeV}$	ATLAS-CONF-2013-061	
	$3^{\text{rd}}$ gen. squarks direct production	$\tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow b\bar{b}$	0	2 b	Yes	20.1	$m(\tilde{b}_1) < 90 \text{ GeV}$	1308.2631
		$\tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow t\bar{t}$	$2 e, \mu$ (SS)	0-3 b	Yes	20.7	$m(\tilde{b}_1) = 2 m(\tilde{t}_1)$	ATLAS-CONF-2013-007
		$\tilde{t}_1 \tilde{t}_1$ (light), $\tilde{t}_1 \rightarrow W b \tilde{\chi}_1^0$	$1-2 e, \mu$	$1-2 b$	Yes	4.7	$m(\tilde{t}_1) = 55 \text{ GeV}$	1208.4905, 1209.2102
		$\tilde{t}_1 \tilde{t}_1$ (medium), $\tilde{t}_1 \rightarrow t \tilde{\chi}_1^0$	$2 e, \mu$	0-2 jets	Yes	20.3	$m(\tilde{t}_1) = m(\tilde{t}_1) + m(W) - 50 \text{ GeV}, m(\tilde{t}_1) < m(\tilde{t}_1)$	1403.4853
		$\tilde{t}_1 \tilde{t}_1$ (medium), $\tilde{t}_1 \rightarrow t \tilde{\chi}_1^0$	$2 e, \mu$	2 jets	Yes	20.3	$m(\tilde{t}_1) = 1 \text{ GeV}$	1403.4853
		$\tilde{t}_1 \tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow t \tilde{\chi}_1^0$	$1 e, \mu$	1 b	Yes	20.1	$m(\tilde{t}_1) < 200 \text{ GeV}, m(\tilde{t}_1) - m(\tilde{t}_1) = 5 \text{ GeV}$	1308.2631
$\tilde{t}_1 \tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow t \tilde{\chi}_1^0$		$1 e, \mu$	1 b	Yes	20.7	$m(\tilde{t}_1) = 0 \text{ GeV}$	ATLAS-CONF-2013-037	
$\tilde{t}_1 \tilde{t}_1$ (natural GMSB)		0	2 b	Yes	20.5	$m(\tilde{t}_1) < 85 \text{ GeV}$	ATLAS-CONF-2013-024	
$\tilde{t}_1 \tilde{t}_1$ (natural GMSB)		0	mono-jet/c-tag	Yes	20.3	$m(\tilde{t}_1) > 150 \text{ GeV}$	ATLAS-CONF-2013-068	
$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow c\bar{c}$		$2 e, \mu$ (Z)	1 b	Yes	20.3	$m(\tilde{t}_1) > 150 \text{ GeV}$	1403.5222	
EW direct	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow t\bar{t}$	$2 e, \mu$	0	Yes	20.3	$m(\tilde{t}_1) < 90 \text{ GeV}$	1403.5294	
	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow t\bar{t} \nu$	$2 e, \mu$	0	Yes	20.3	$m(\tilde{t}_1) = 0 \text{ GeV}, m(\tilde{t}_1, \nu) = 0.5(m(\tilde{t}_1) + m(\tilde{t}_1))$	1403.5294	
	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow t\bar{t} \nu$	$2 e, \mu$	0	Yes	20.7	$m(\tilde{t}_1) = 0 \text{ GeV}, m(\tilde{t}_1, \nu) = 0.5(m(\tilde{t}_1) + m(\tilde{t}_1))$	ATLAS-CONF-2013-028	
	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow t\bar{t} \nu$	$3 e, \mu$	0	Yes	20.3	$m(\tilde{t}_1) = 0 \text{ GeV}, m(\tilde{t}_1, \nu) = 0.5(m(\tilde{t}_1) + m(\tilde{t}_1))$	1402.7029	
	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow t\bar{t} \nu$	$2 e, \mu$	0	Yes	20.3	$m(\tilde{t}_1) = m(\tilde{t}_1), m(\tilde{t}_1) = 0, m(\tilde{t}_1) = 0, \text{ sleptons decoupled}$	1403.5294, 1402.7029	
	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow t\bar{t} \nu$	$1 e, \mu$	2 b	Yes	20.3	$m(\tilde{t}_1) = m(\tilde{t}_1), m(\tilde{t}_1) = 0, \text{ sleptons decoupled}$	ATLAS-CONF-2013-093	
	Long-lived particles	Direct $\tilde{t}_1 \tilde{t}_1$ prod., long-lived $\tilde{t}_1$	Disapp. trk	1 jet	Yes	20.3	$m(\tilde{t}_1) = m(\tilde{t}_1) = 160 \text{ MeV}, \tau(\tilde{t}_1) = 0.2 \text{ ns}$	ATLAS-CONF-2013-089
		Stable, stopped $\tilde{g}$ R-hadron	0	1-5 jets	Yes	22.9	$m(\tilde{g}) = 100 \text{ GeV}, 10 \mu\text{s} < \tau(\tilde{g}) < 1000 \text{ s}$	ATLAS-CONF-2013-057
		GMSB, stable $\tilde{g}$ , $\tilde{g} \rightarrow c\bar{c} \nu, \beta \rightarrow \tau \nu, \epsilon \rightarrow \mu \nu$	$1-2 \mu$	-	-	15.9	$10^{-10} \text{ s} < \tau < 10^{-9} \text{ s}$	ATLAS-CONF-2013-058
		GMSB, $\tilde{g} \rightarrow c\bar{c} \nu$ , long-lived $\tilde{g}$	$2 \gamma$	-	Yes	4.7	$0.4 \cdot \tau(\tilde{g}) > 2 \text{ ns}$	1304.6310
$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}$ (RPV)		$1 \mu$ , displ. vtx	-	-	20.3	$1.5 \cdot c\tau < 156 \text{ mm}, \text{BR}(\mu) = 1, m(\tilde{g}) = 108 \text{ GeV}$	ATLAS-CONF-2013-092	
RPV		LFV $pp \rightarrow \tilde{g}\tilde{g} + X, \tilde{g}\tilde{g} \rightarrow e\mu$	$2 e, \mu$	-	-	4.6	$A_{11} = -0.10, A_{133} = 0.05$	1212.1272
		LFV $pp \rightarrow \tilde{g}\tilde{g} + X, \tilde{g}\tilde{g} \rightarrow e\mu + \tau$	$1 e, \mu + \tau$	-	-	4.6	$A_{11} = -0.10, A_{133} = 0.05$	1212.1272
		Bilinear RPV CMSSM	$1 e, \mu$	7 jets	Yes	4.7	$m(\tilde{g}) = 0 \text{ GeV}, c\tau_{\tilde{g}} > 1 \text{ mm}$	ATLAS-CONF-2012-140
		$\tilde{g}\tilde{g}, \tilde{g} \rightarrow c\bar{c} \nu, \beta \rightarrow \tau \nu, \epsilon \rightarrow \mu \nu$	$4 e, \mu$	-	Yes	20.7	$m(\tilde{g}) > 300 \text{ GeV}, A_{133} > 0$	ATLAS-CONF-2013-036
		$\tilde{g}\tilde{g}, \tilde{g} \rightarrow c\bar{c} \nu, \beta \rightarrow \tau \nu, \epsilon \rightarrow \mu \nu$	$3 e, \mu + \tau$	-	Yes	20.7	$m(\tilde{g}) > 80 \text{ GeV}, A_{133} > 0$	ATLAS-CONF-2013-036
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}$	0	6-7 jets	-	20.3	$\text{BR}(\tilde{g}) = \text{BR}(\tilde{g}) - \text{BR}(\tilde{g}) = 0\%$	ATLAS-CONF-2013-091	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\bar{b}$	$2 e, \mu$ (SS)	0-3 b	Yes	20.7		ATLAS-CONF-2013-007	
	Other	Scalar gluon pair, sgluon $\rightarrow q\bar{q}$	0	4 jets	-	4.6	incl. limit from 1110.2693	1210.4826
		Scalar gluon pair, sgluon $\rightarrow t\bar{t}$	$2 e, \mu$ (SS)	2 b	Yes	14.3		ATLAS-CONF-2013-051
		WIMP interaction (D5, Dirac $\chi$ )	0	mono-jet	Yes	10.5	$m(\chi) > 80 \text{ GeV}$ , limit of $\sim 687 \text{ GeV}$ for D8	ATLAS-CONF-2012-147

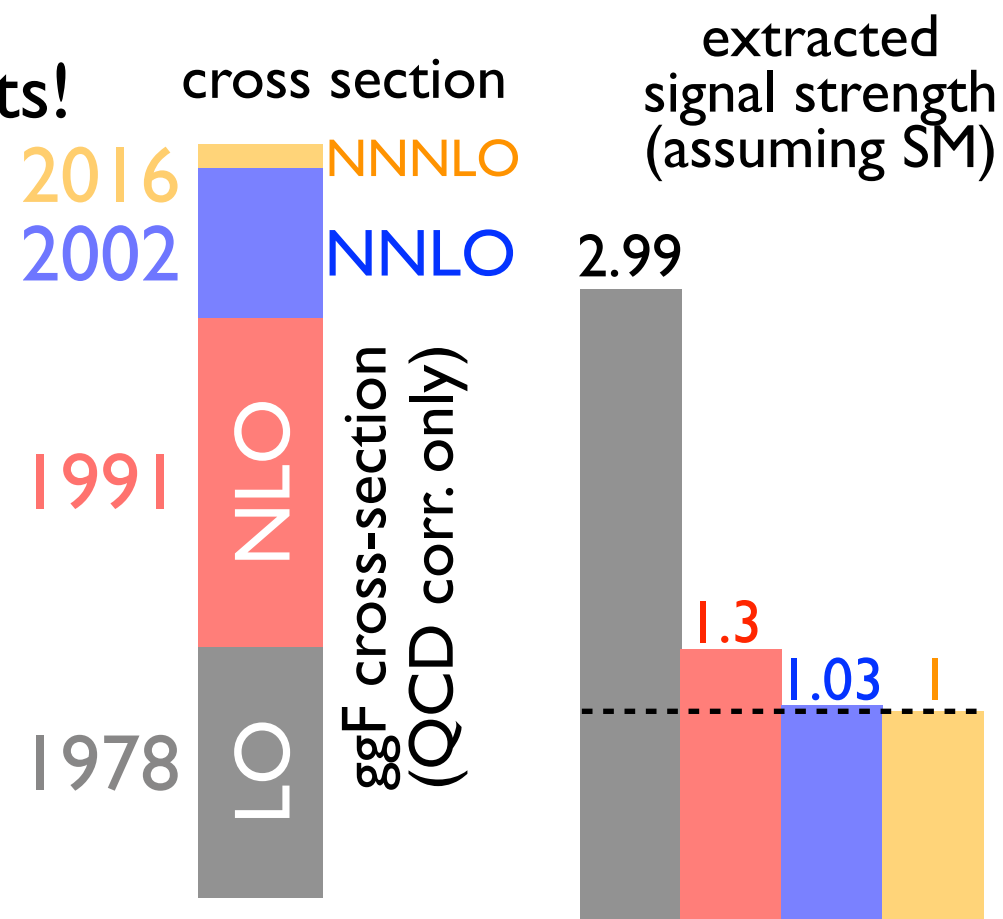
\*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus  $1\sigma$  theoretical signal cross section uncertainty.

# Cross-section measurements

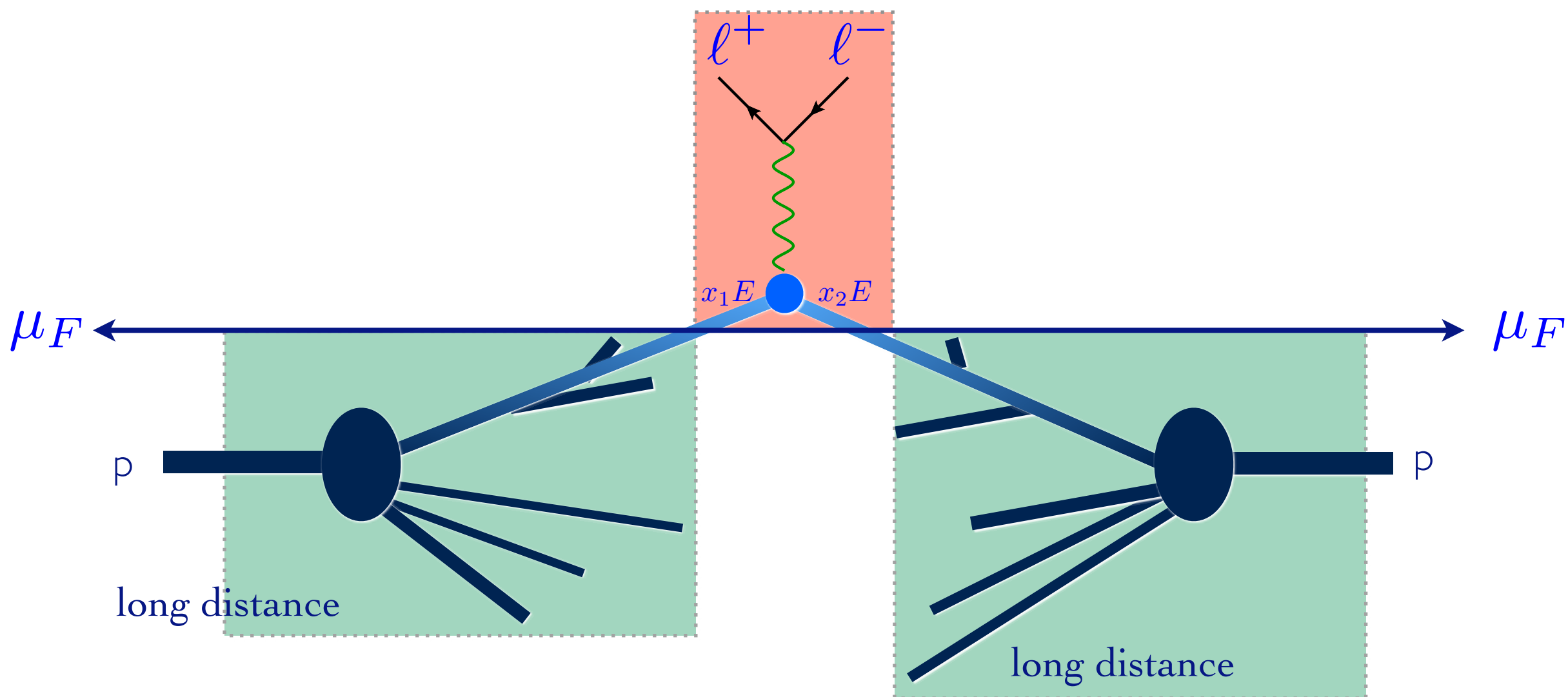
- The discovery of the Higgs boson is an emblematic example of the need for precision
- Large perturbative corrections for the dominant channel (gluon fusion)
- Without higher-order corrections, measured signal strength  $\sim 3 * SM$
- Very competitive experimental measurements!



$$\mu = \frac{\sigma_{\text{EXP}}}{\sigma_{\text{TH}}}$$



# How to compute a cross-section



$$\sum_{a,b} \int dx_1 dx_2 d\Phi_{\text{FS}} f_a(x_1, \mu_F) f_b(x_2, \mu_F) \hat{\sigma}_{ab \rightarrow X}(\hat{S}, \mu_F, \mu_R)$$

Phase-space integral
Parton density functions
Parton-level cross section





# Perturbation theory at work

$$\hat{\sigma}_{ab \rightarrow X}(\hat{s}, \mu_F, \mu_R) \quad \text{Parton-level cross section}$$

The parton-level cross section can be computed as a series in perturbation theory, using the coupling constant as an expansion parameter

$$\hat{\sigma} = \alpha_s^b \sigma_0 + \alpha_s^{b+1} \sigma_1 + \alpha_s^{b+2} \sigma_2 + \alpha_s^{b+3} \sigma_3 + \dots$$

Remember:

$$\alpha_s = \alpha_s(\mu_R) \quad \sigma_i = \sigma_i(\mu_R, \mu_F)$$

Coupling and cross section depend on *unphysical* scales

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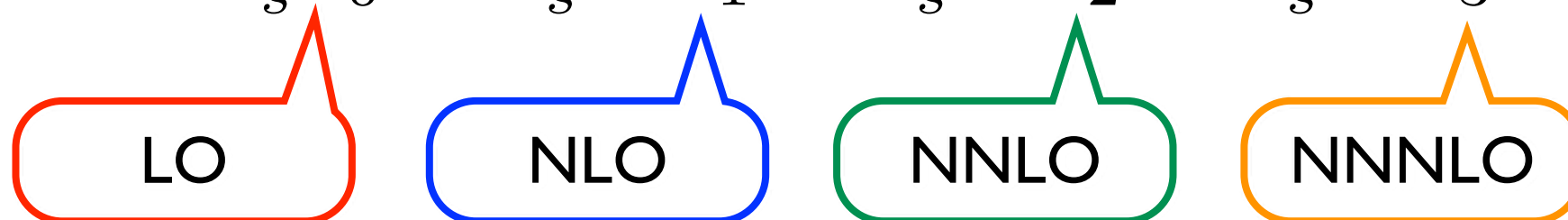
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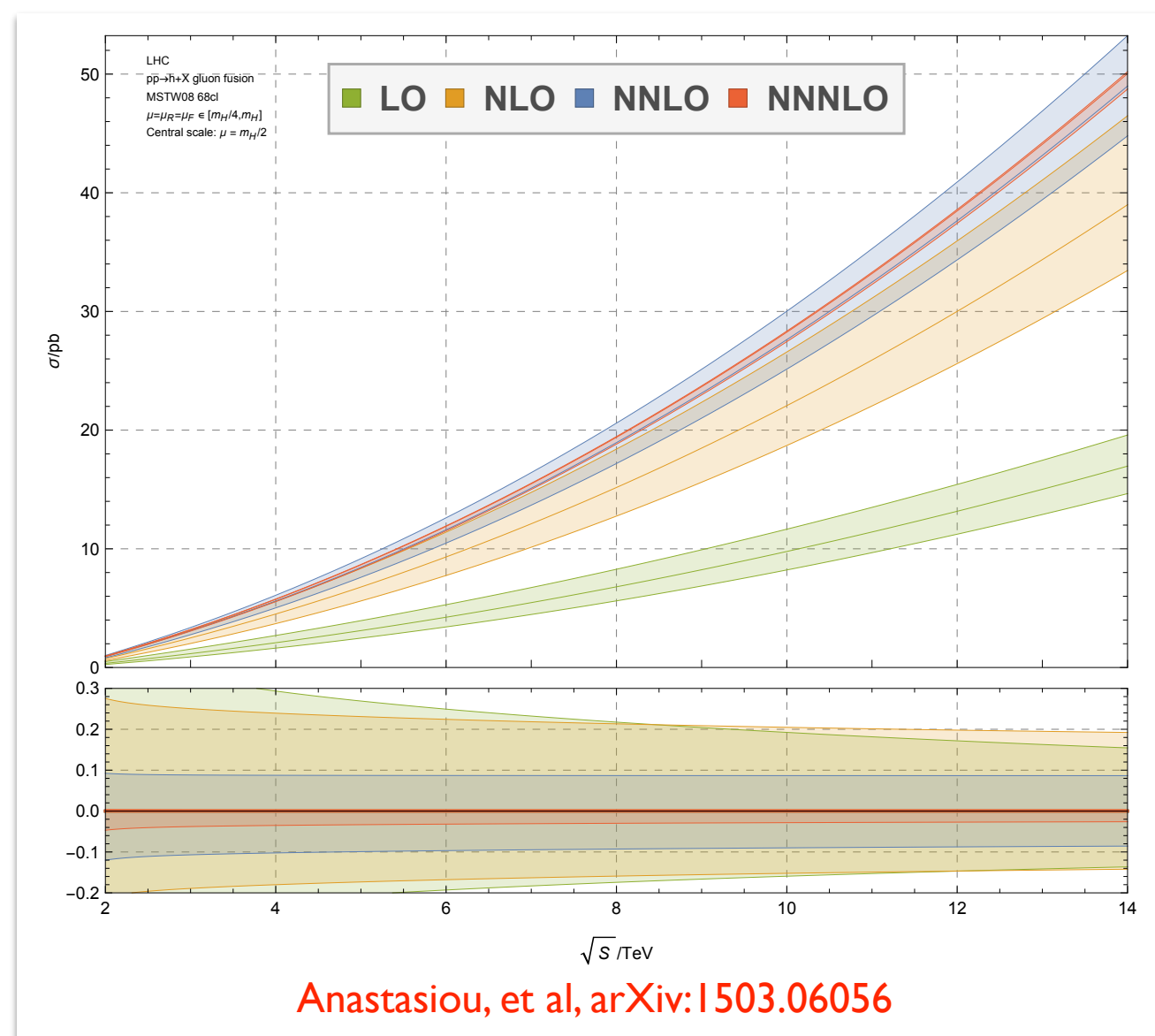
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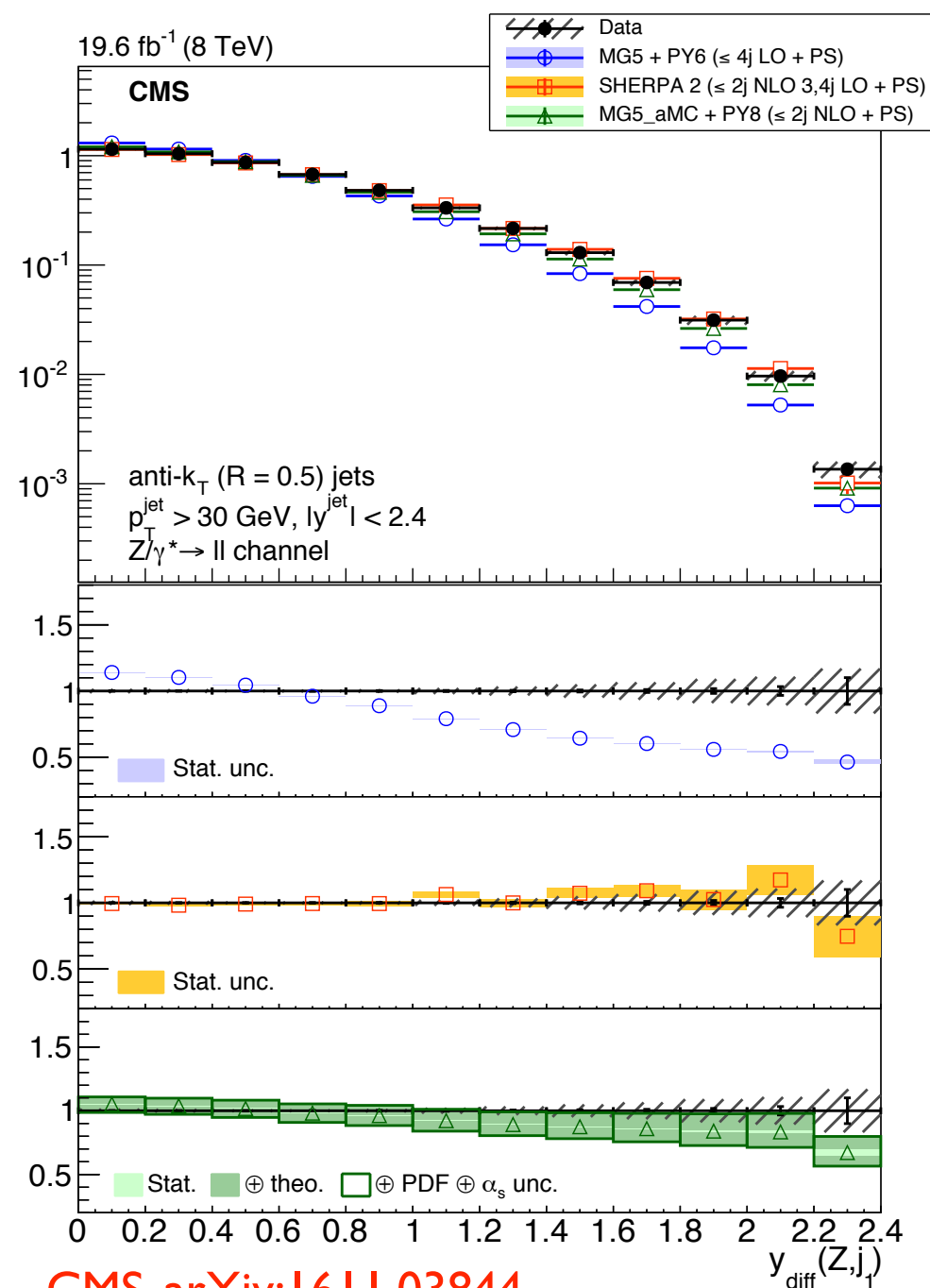


# Perturbation theory at work

- The inclusion of higher orders improves the reliability of a given computation
  - More reliable description of total rates and shapes
  - Residual uncertainties related to the arbitrary scales in the process decrease
  - The computational complexity grows exponentially
  - NLO is mandatory for LHC physics!



# Perturbation theory at work



CMS, arXiv:1611.03844

- In order to describe data, LO predictions must be rescaled to match the cross section including higher orders (typically NNLO)
- NLO predictions are generally not rescaled → More predictive power
- NLO effects can be important even if merged samples are used at LO

## In these lectures:

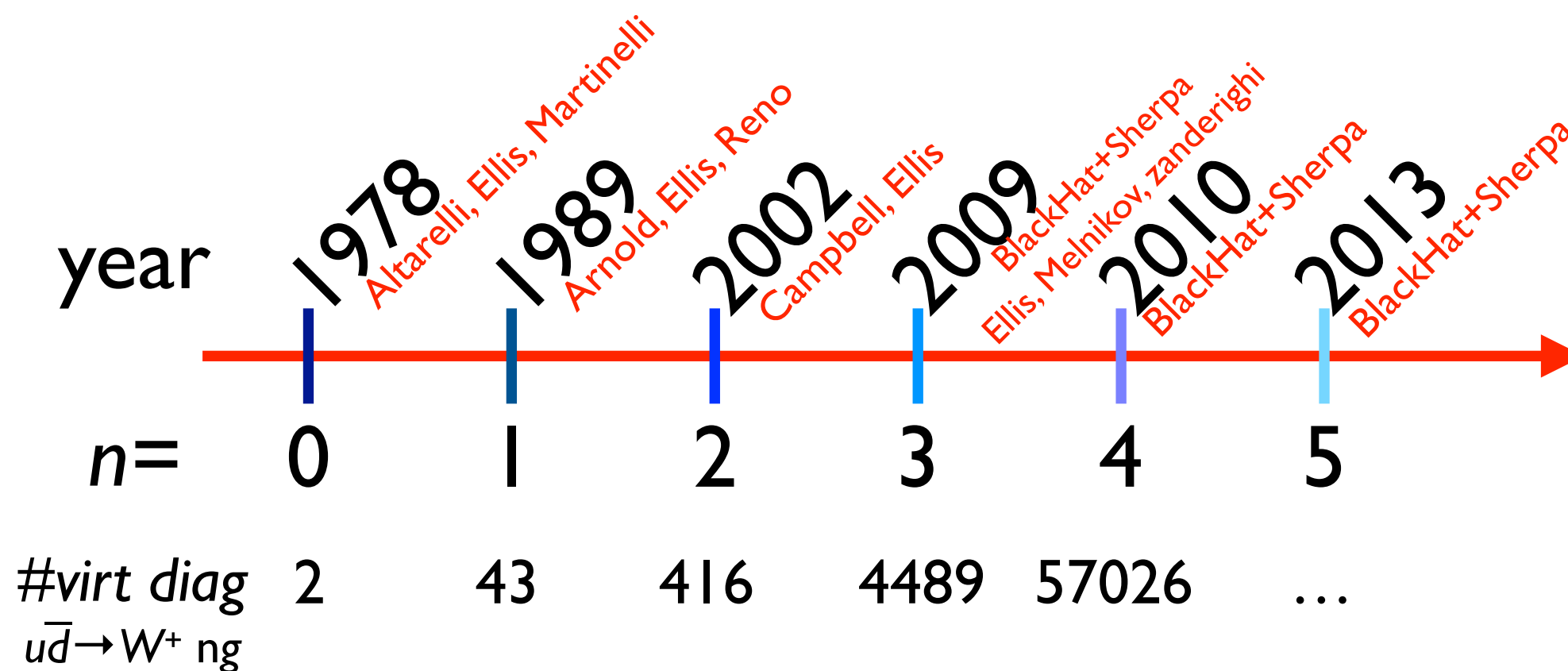
- How to compute effectively a NLO cross section?
- How to deal with infrared divergences?
- ~~How to compute loops?~~
- How about EW corrections?



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**#141 171**

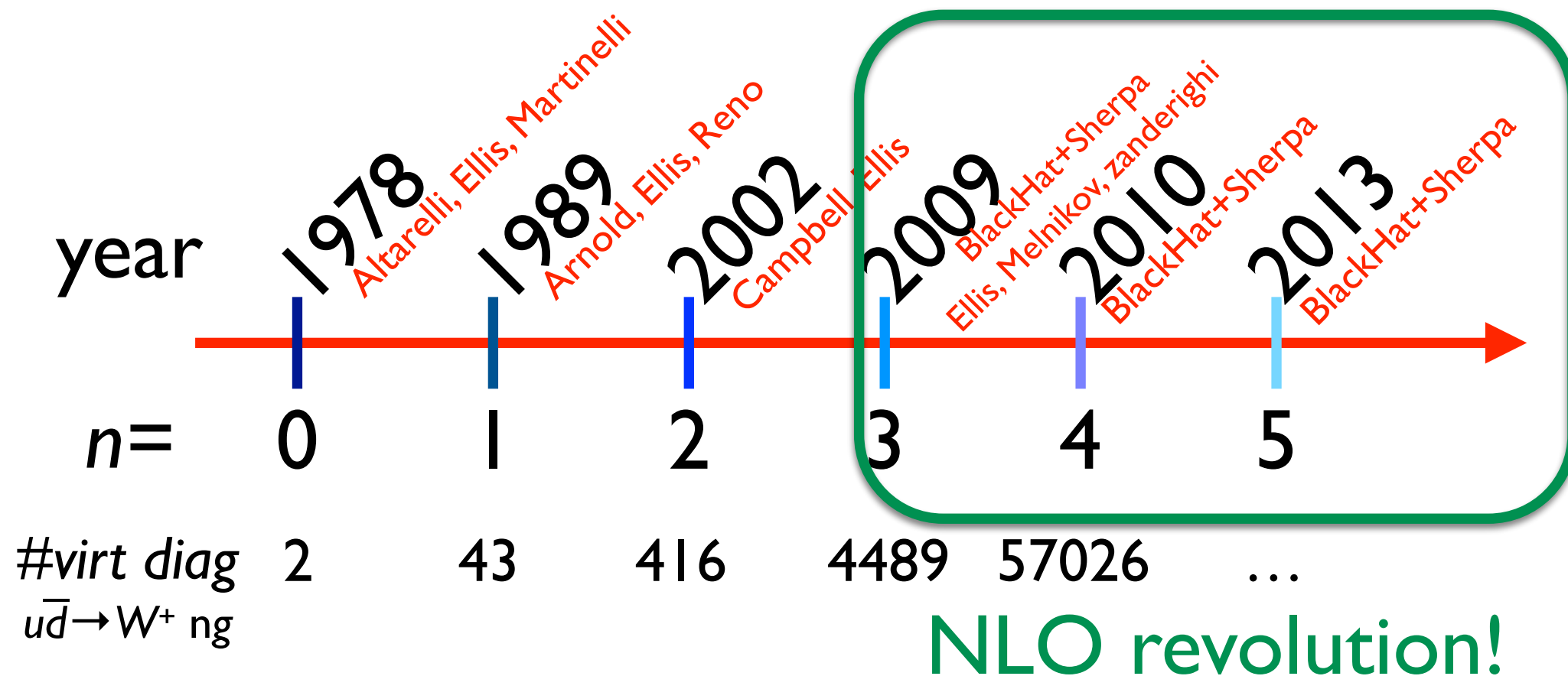
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- NLO evolution:
  - e.g.  $pp \rightarrow W+n$  jets



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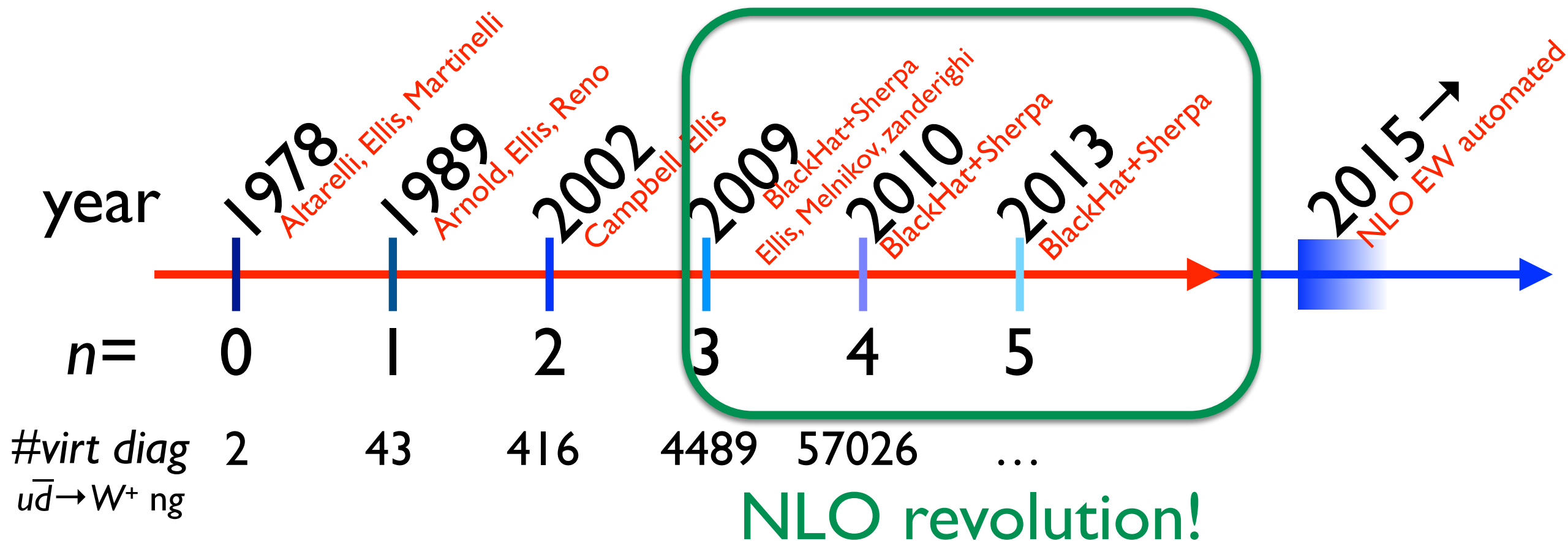
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# NLO (pre)history

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# NLO revolution

- Amazing development of computational techniques to tackle *any* process at NLO

- Local subtraction

Frixione, Kunszt, Signer, hep-ph/9512328  
Catani, Seymour, hep-ph/9605323

- Computation of loop MEs

- Tensor reduction

Passarino, Veltman, 1979

Denner, Dittmaier, hep-ph/509141

Binoth, Guillet, Heinrich, Pilon, Reiter, arXiv:0810.0992

- Generalized unitarity

Bern, Dixon, Dunbar, Kosower, hep-ph/9403226 + ...

Ellis, Giele, Kunszt, arXiv:0708.2398

+ Melnikov, arXiv:0806.3467

- Integrand reduction

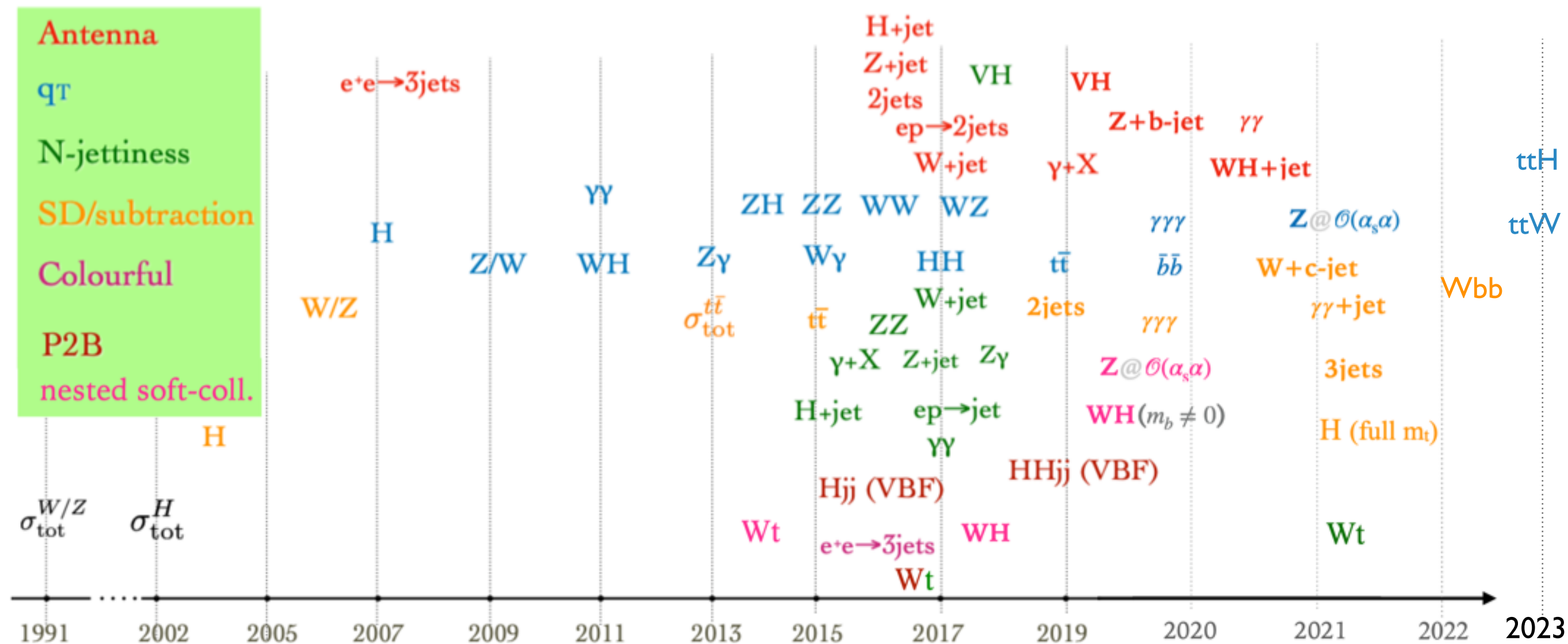
Ossola, Papadopoulos, Pittau, hep-ph/0609007

Del Aguila, Pittau, hep-ph/0404120

Mastrolia, Ossola, Reiter, Tramontano, arXiv:1006.0710



# The NNLO revolution is happening now!



Adapted from G. Zanderighi @LHCP23



# Going NLO

$$\hat{\sigma} = \alpha_s^b \sigma_0 + \alpha_s^{b+1} \sigma_1 + \alpha_s^{b+2} \sigma_2 + \alpha_s^{b+3} \sigma_3 + \dots$$

- NLO is the first order where the scale dependence in  $\alpha_s$  and PDFs is compensated by loop corrections
  - First reliable predictions for rates and uncertainties
- Better description of final state (inclusion of extra radiation)
- Opening of new partonic channels from real emissions
- Learning NLO technicalities will set the basis for us (you!) to tackle NNLO or beyond

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NNLO

NNNLO

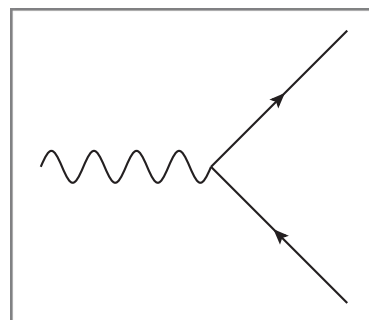
- NLO is the first order where the scale dependence in  $\alpha_s$  and PDFs is compensated by loop corrections
  - First reliable predictions for rates and uncertainties
- Better description of final state (inclusion of extra radiation)
- Opening of new partonic channels from real emissions
- Learning NLO technicalities will set the basis for us (you!) to tackle NNLO or beyond

# NLO: how to?

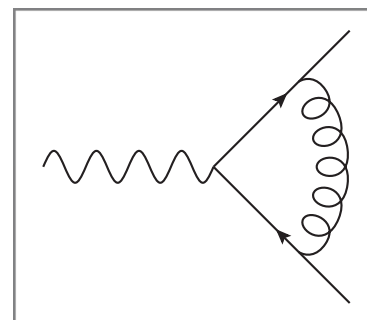
- Three ingredients need to be computed at NLO

$$\sigma_{NLO} = \int_n \alpha_s^b d\sigma_0 + \int_n \alpha_s^{b+1} d\sigma_V + \int_{n+1} \alpha_s^{b+1} d\sigma_R$$

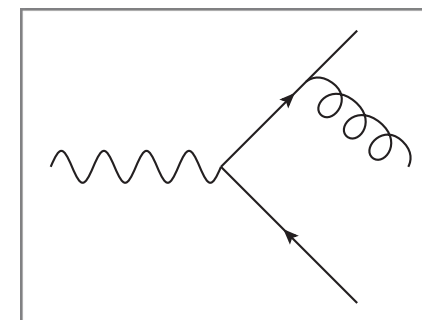
↑  
Born  
cross section



↑  
Virtual  
corrections

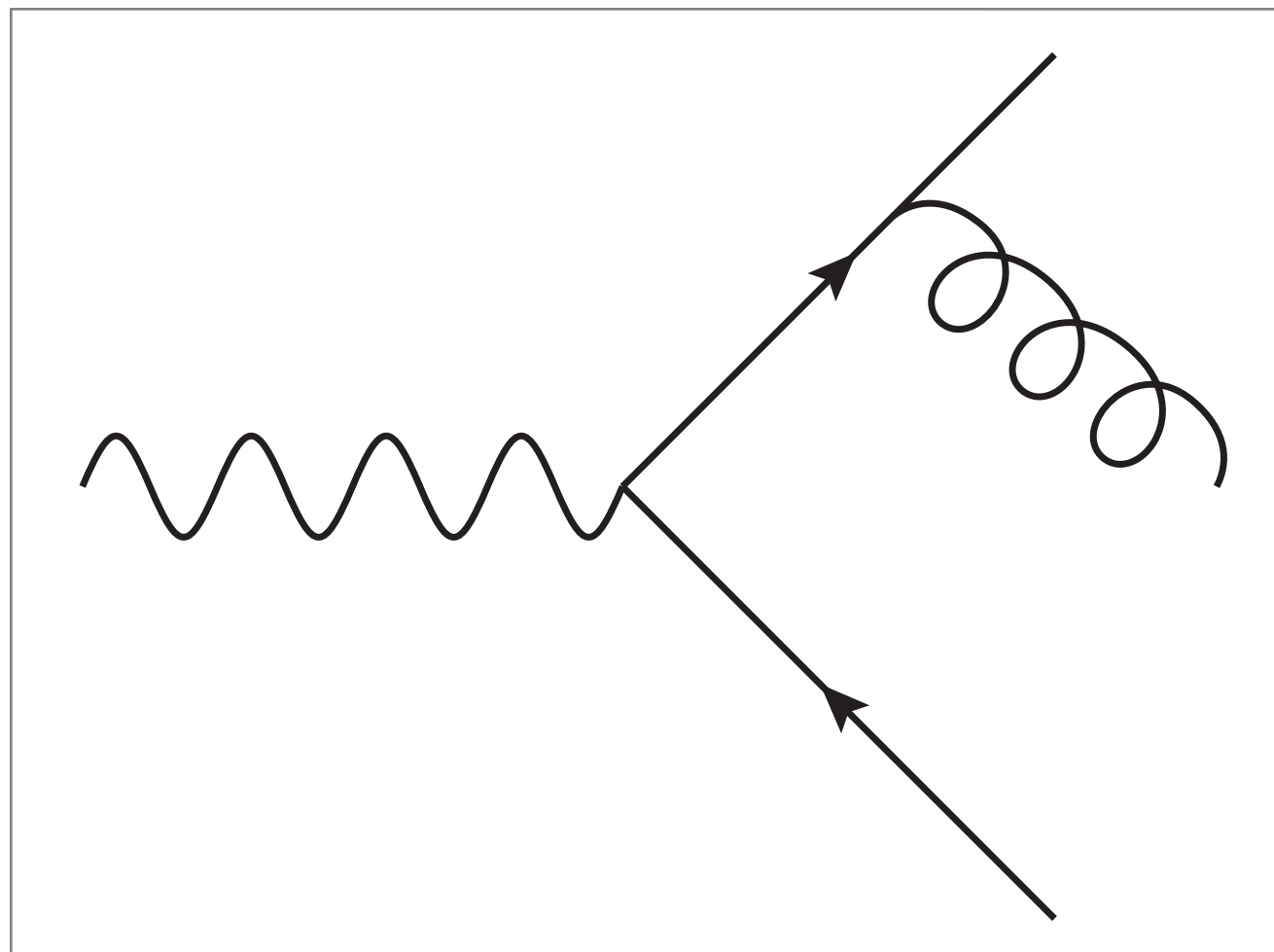


↑  
Real-emission  
corrections



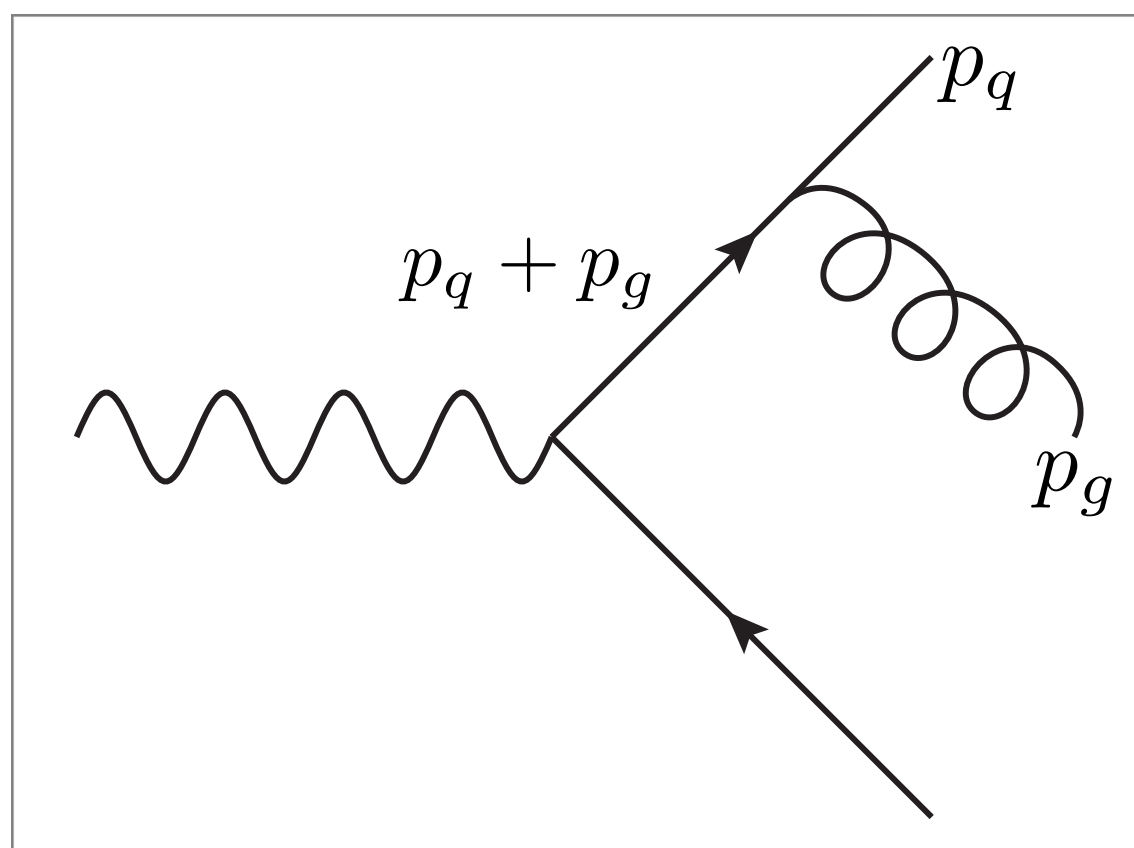
- Remember: virtual and reals are not separately finite, but their sum is (KLN theorem). Divergences have to be subtracted before numerical integration. We will shortly see how

# Infrared divergences



$$\sigma_{NLO} = \int_n \alpha_s^b d\sigma_0 + \int_n \alpha_s^{b+1} d\sigma_V + \int_{\underline{n+1}} \alpha_s^{b+1} d\sigma_R$$

# Branching

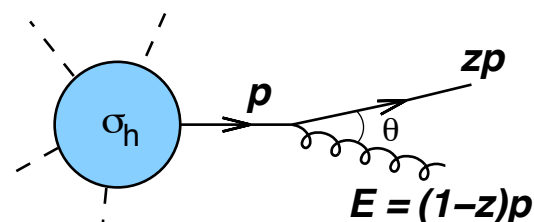


$$\int_{n+1} \alpha_s^{b+1} d\sigma_R$$

- When the integral over the phase-space of the gluon is performed, one can have  $(p_q+p_g)^2=0$
- Since  $(p_q+p_g)^2=2E_qE_g(1-\cos\theta)$  it happens when the gluon is soft ( $E_g=0$ ) or collinear to the quark ( $\theta=0$ )
- In both cases, the propagator leads to a divergent cross section

# Singularities

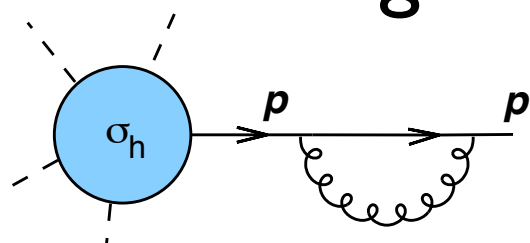
- Let us rewrite the branching of a gluon from a quark as



$$\sigma_{h+g} \simeq \sigma_h \frac{\alpha_s C_F}{\pi} \frac{dz}{1-z} \frac{dk_t^2}{k_t^2}$$

Where  $k_t$  is the transverse momentum of the gluon  $k_t = E \sin \theta$ .  
It diverges in the soft ( $z \rightarrow 1$ ) and collinear ( $k_t \rightarrow 0$ ) region

- These singularities cancel with the virtual contribution, which comes from the integration of the loop momentum



$$\sigma_{h+V} \simeq -\sigma_h \frac{\alpha_s C_F}{\pi} \frac{dz}{1-z} \frac{dk_t^2}{k_t^2}$$

- The cancelation happens if we cannot distinguish between the case of no branching, and that of a soft/collinear branching





# Cancellation of divergences

- The KLN theorem tells us that divergences from the virtual and real emission cancel in the sum *if observables are insensitive to soft and collinear branchings* (IR-safety)
- When doing an analytic computation in dimensional regularisation, divergences appear as poles in the regularisation parameter  $\varepsilon$
- In the real emissions, poles appear *after* the phase space integration in  $d$  dimension



# Infrared safety

- In order to have meaningful predictions in fixed-order perturbation theory, observables must be IR-safe, *i.e.* not sensitive to the emission of soft or collinear partons.
- In particular, if an observable depends on the momentum  $p_i$ , it must not be sensitive on the branching  $p_i \rightarrow p_j + p_k$ , where either  $p_j$  is soft or  $p_j$  and  $p_k$  are collinear
- For example
  - The number of gluons in an event
  - The number of jets with  $p_T > p_T^{min}$
  - The hardest parton in an event
  - The hardest jet

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- For example
  - The number of gluons in an event is not IR-safe ❌
  - The number of jets with  $p_T > p_T^{min}$  is IR-safe ✅
  - The hardest parton in an event is not IR-safe ❌
  - The hardest jet is IR-safe ✅



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# Phase space integration

$$\sigma_{NLO} = \int d^4\Phi_n \mathcal{B} + \int d^4\Phi_n \mathcal{V} + \int d^4\Phi_{n+1} \mathcal{R}$$

contains  $\int d^d l$

- For complicated processes the integrations have to be done via MonteCarlo techniques, in an integer number of dimensions
- Divergences have to be canceled explicitly
- Slicing/Subtraction methods have been developed to extract divergences from the phase-space integrals

# Example

- Suppose that we can cast the phase space integral in the form

$$\int_0^1 dx f(x) \quad \text{with} \quad f(x) = \frac{g(x)}{x} \quad \text{and} \quad g(x) \text{ a regular function}$$

- We introduce a regulator which renders the integral finite

$$\int_0^1 dx x^\varepsilon f(x) = \int_0^1 dx \frac{g(x)}{x^{1-\varepsilon}}$$

- The divergence will turn into a pole in  $\varepsilon$ . How can we extract the pole?

# Phase space slicing

$$\lim_{\varepsilon \rightarrow 0} \int_0^1 dx x^\varepsilon f(x) = \lim_{\varepsilon \rightarrow 0} \int_0^1 dx \frac{g(x)}{x^{1-\varepsilon}}$$

- We introduce a small parameter  $\delta \ll 1$ :

$$\lim_{\varepsilon \rightarrow 0} \int_0^1 dx \frac{g(x)}{x^{1-\varepsilon}} = \lim_{\varepsilon \rightarrow 0} \left( \int_0^\delta dx \frac{g(x)}{x^{1-\varepsilon}} + \int_\delta^1 dx \frac{g(x)}{x^{1-\varepsilon}} \right)$$

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**pole in  $\varepsilon$**

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$$\simeq \lim_{\varepsilon \rightarrow 0} \left( \int_0^\delta dx \frac{g(0)}{x^{1-\varepsilon}} + \int_\delta^1 dx \frac{g(x)}{x^{1-\varepsilon}} \right)$$

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pole in  $\varepsilon$

finite integral

(can be computed numerically)



# Subtraction method

$$\lim_{\varepsilon \rightarrow 0} \int_0^1 dx x^\varepsilon f(x) = \lim_{\varepsilon \rightarrow 0} \int_0^1 dx \frac{g(x)}{x^{1-\varepsilon}}$$

- Add and subtract  $g(0)/x$

$$\lim_{\varepsilon \rightarrow 0} \int_0^1 dx \frac{g(x)}{x^{1-\varepsilon}} = \lim_{\varepsilon \rightarrow 0} \int_0^1 dx x^\varepsilon \left( \frac{g(0)}{x} + \frac{g(x)}{x} - \frac{g(0)}{x} \right)$$



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$$= \lim_{\varepsilon \rightarrow 0} \boxed{\frac{1}{\varepsilon} g(0)} + \int_0^1 dx \boxed{\frac{g(x) - g(0)}{x}} \quad \text{finite integral}$$

**pole in  $\varepsilon$**  (can be computed numerically)

# Slicing vs Subtraction

- In both cases the pole is extracted and we end up with a finite remainder:

$$g(0) \log \delta + \int_{\delta}^1 dx \frac{g(x)}{x}$$

$$\int_0^1 dx \frac{g(x) - g(0)}{x}$$

- Subtraction acts like a plus distribution
- Slicing works only for small  $\delta$ :  $\delta$ -independence of cross section and distributions must be proven; subtraction is exact
- Both methods have cancelations between large numbers. If for a given observable  $\lim_{x \rightarrow 0} O(x) \neq O(0)$  or we choose a too small bin size, instabilities will arise (we cannot ask for an infinite resolution)
- Subtraction is in general more flexible: good for automation

# NLO with subtraction

$$\sigma_{NLO} = \int d^4\Phi_n \mathcal{B} + \int d^4\Phi_n \mathcal{V} + \int d^4\Phi_{n+1} \mathcal{R}$$

- With the subtraction terms the expression becomes

$$\begin{aligned} \sigma_{NLO} = & \int d^4\Phi_n \mathcal{B} \\ & + \int d^4\Phi_n \left( \mathcal{V} + \int d^d\Phi_1 \mathcal{C} \right)_{\varepsilon \rightarrow 0} \\ & + \int d^4\Phi_{n+1} (\mathcal{R} - \mathcal{C}) \end{aligned}$$

- Terms in brackets are finite and can be integrated numerically in  $d=4$  and independently one from another

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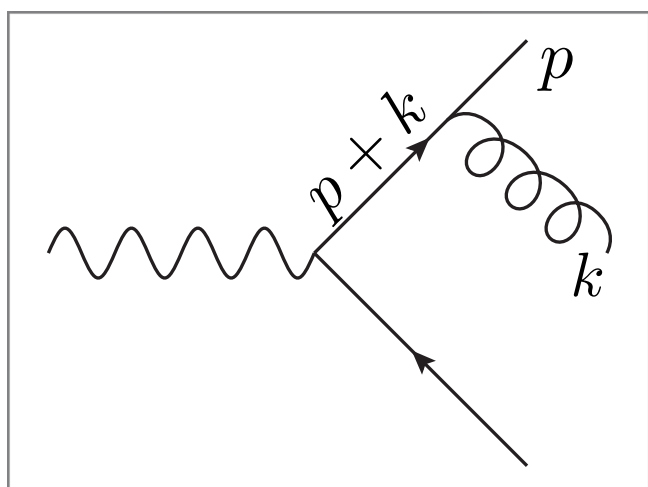


# The subtraction term

- The subtraction term  $C$  should be chosen such that:
  - It exactly matches the singular behaviour of  $R$
  - It can be integrated numerically in a convenient way
  - It can be integrated exactly in  $d$  dimension, leading to the soft and/or collinear poles in the dimensional regulator
  - It is process independent (overall factor times Born)

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  - It can be integrated exactly in  $d$  dimension, leading to the soft and/or collinear poles in the dimensional regulator
  - It is process independent (overall factor times Born)
- QCD comes to help: structure of divergences is universal:



$$(p+k)^2 = 2E_p E_k (1 - \cos \theta_{pk})$$

- Collinear singularity:

$$\lim_{p//k} |M_{n+1}|^2 \simeq |M_n|^2 P^{AP}(z)$$

- Soft singularity:

$$\lim_{k \rightarrow 0} |M_{n+1}|^2 \simeq \sum_{ij} |M_n^{ij}|^2 \frac{p_i p_j}{p_i k p_j k}$$





# Two subtraction methods

## Dipole subtraction

Catani, Seymour, [hep-ph/9602277](#) & [hep-ph/9605323](#)

- Recoil taken by one parton  
→  $N^3$  scaling
- Method evolves from cancelation of soft divergences
- Proven to work for simple and complicated processes
- Automated in MadDipole, AutoDipole, Sherpa, Helac-NLO, ...

## FKS subtraction

Frixione, Kunszt, Signer, [hep-ph/9512328](#)

- Recoil distributed among all particles  
→  $N^2$  scaling
- Method evolves from cancelation of collinear divergences
- Proven to work for simple and complicated processes
- Automated in MadGraph5\_aMC@NLO and in the Powheg box/Powhel

# FKS subtraction #1

## Phase space partition

- Let us consider the real emission

$$d\sigma_R = |M^{n+1}|^2 d\Phi_{n+1}$$

- The matrix element  $|M^{n+1}|^2$  diverges as

$$|M^{n+1}| \sim \frac{1}{\xi_i^2} \frac{1}{1 - y_{ij}}$$

$$\xi_i = E_i \sqrt{\hat{s}}$$

$$y_{ij} = \cos \theta_{ij}$$

- Partition the phase space in order to have at most one soft and one collinear singularity

$$d\sigma_R = \sum_{ij} S_{ij} |M^{n+1}|^2 d\Phi_{n+1} \quad \sum_{ij} S_{ij} = 1$$

$$S_{ij} \rightarrow 1 \text{ if } k_i \cdot k_j \rightarrow 0$$

$$S_{ij} \rightarrow 0 \text{ if } k_{m \neq i} \cdot k_{n \neq j} \rightarrow 0$$

# FKS subtraction #2

## Plus prescriptions

- Use plus prescriptions in  $y_{ij}$  and  $\xi_i$  to subtract the divergences

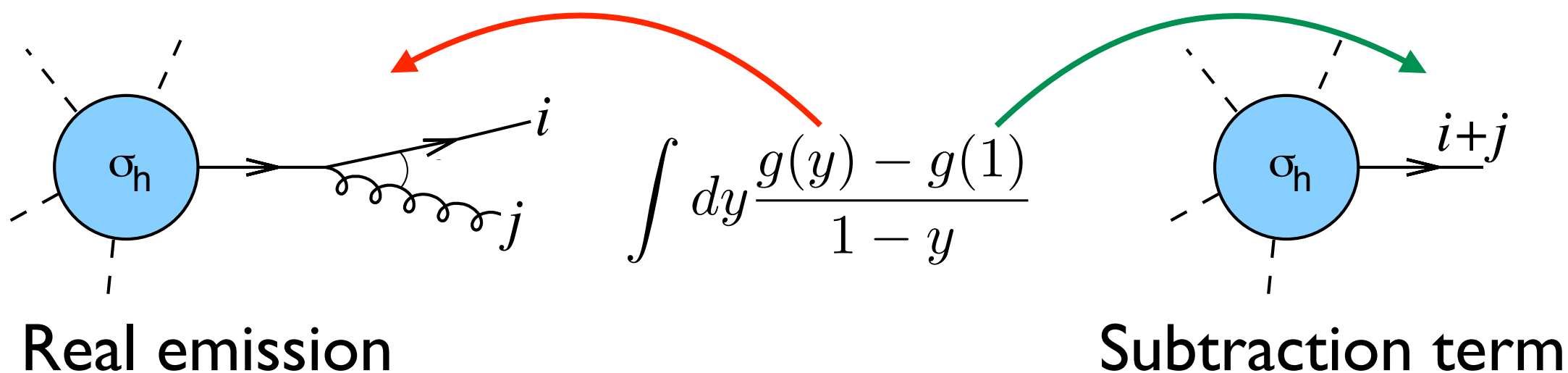
$$d\sigma_{\tilde{R}} = \sum_{ij} \left( \frac{1}{\xi_i} \right)_+ \left( \frac{1}{1 - y_{ij}} \right)_+ \xi_i (1 - y_{ij}) S_{ij} |M^{n+1}|^2 d\Phi_{n+1}$$

- Plus prescriptions are defined as

$$\int d\xi \left( \frac{1}{\xi} \right)_+ f(\xi) = \int d\xi \frac{f(\xi) - f(0)}{\xi} \quad \int dy \left( \frac{1}{1 - y} \right)_+ g(y) = \int dy \frac{g(y) - g(1)}{1 - y}$$

- Maximally three counterevents are needed
  - Soft counterevent ( $\xi_i \rightarrow 0$ )
  - Collinear counterevents ( $y_{ij} \rightarrow 1$ )
  - Soft-collinear counterevents ( $\xi_i \rightarrow 0$  and  $y_{ij} \rightarrow 1$ )
- The counterevents will feature the *same* kinematics

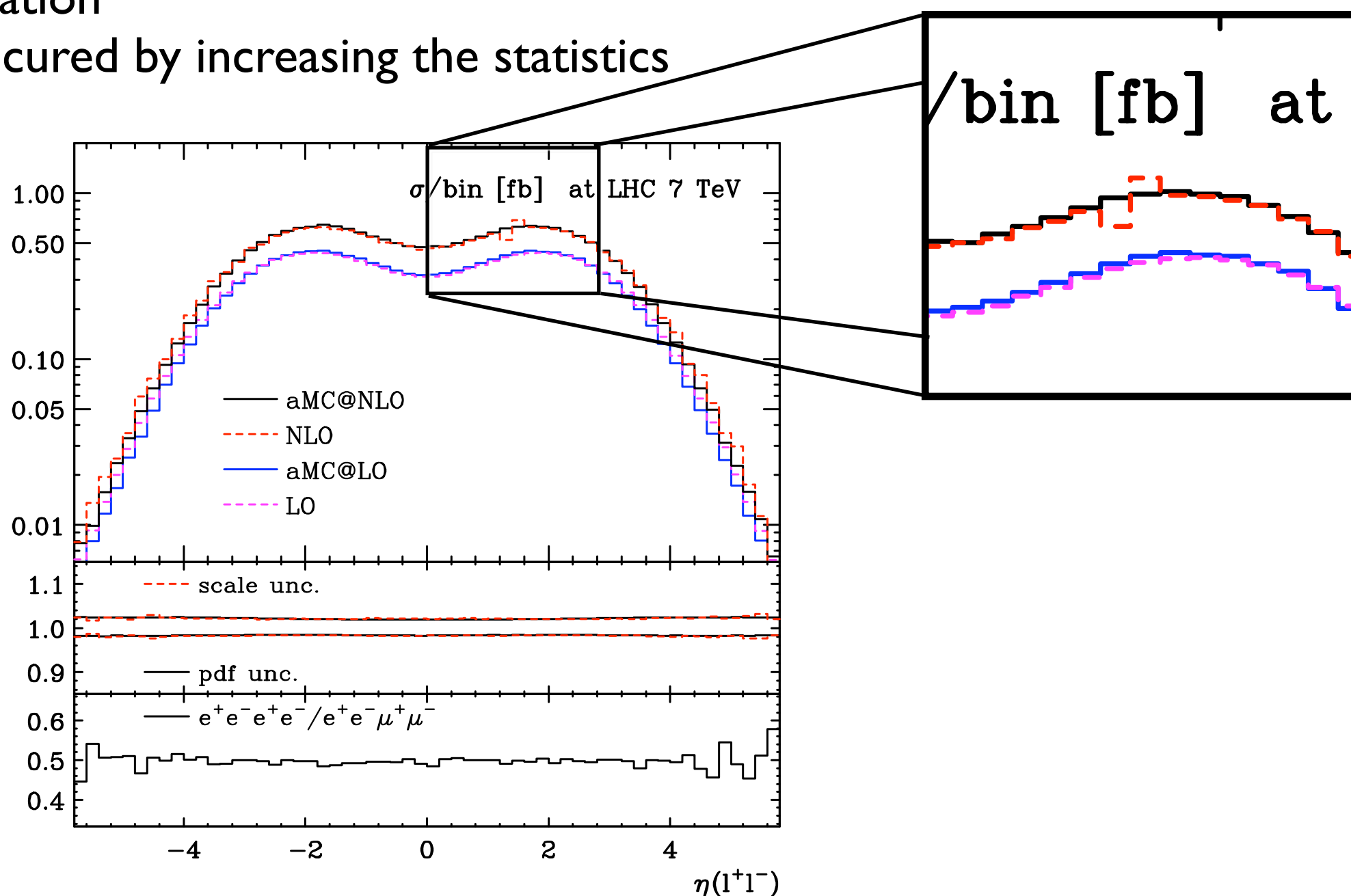
# Kinematics of counterevents



- If  $i$  and  $j$  are on-shell in the event, for the counterevent the combined particle  $i+j$  must be on shell
- $i+j$  can be put on shell only by reshuffling the momenta of the other particles
- It can happen that event and counterevent end up in different histogram bins
  - Use IR-safe observables and don't ask for infinite resolution!
  - Still, these precautions do not eliminate the problem...

# An example in 4-lepton production

- The NLO result shows the typical peak-dip structure that hampers fixed-order computation
- Can be cured by increasing the statistics





# Can we generate unweighted events at NLO?

- Another consequence of the kinematic mismatch is that we cannot generate events at NLO
- $n+1$ -body contribution and  $n$ -body contribution are not bounded from above  $\rightarrow$  unweighting not possible
- Further ambiguity on which kinematics to use for the unweighted events



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More tomorrow

# Filling histograms on-the-fly

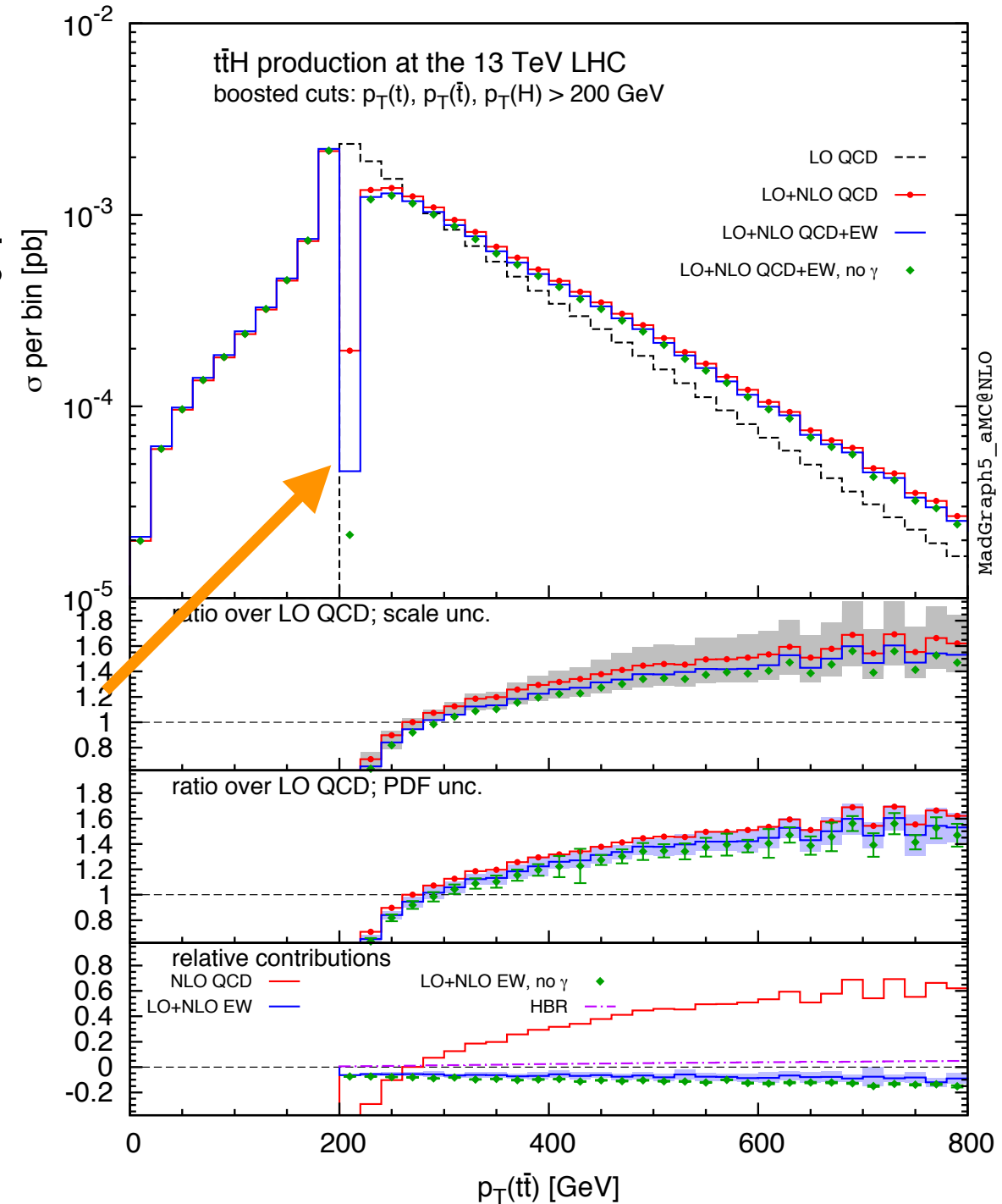
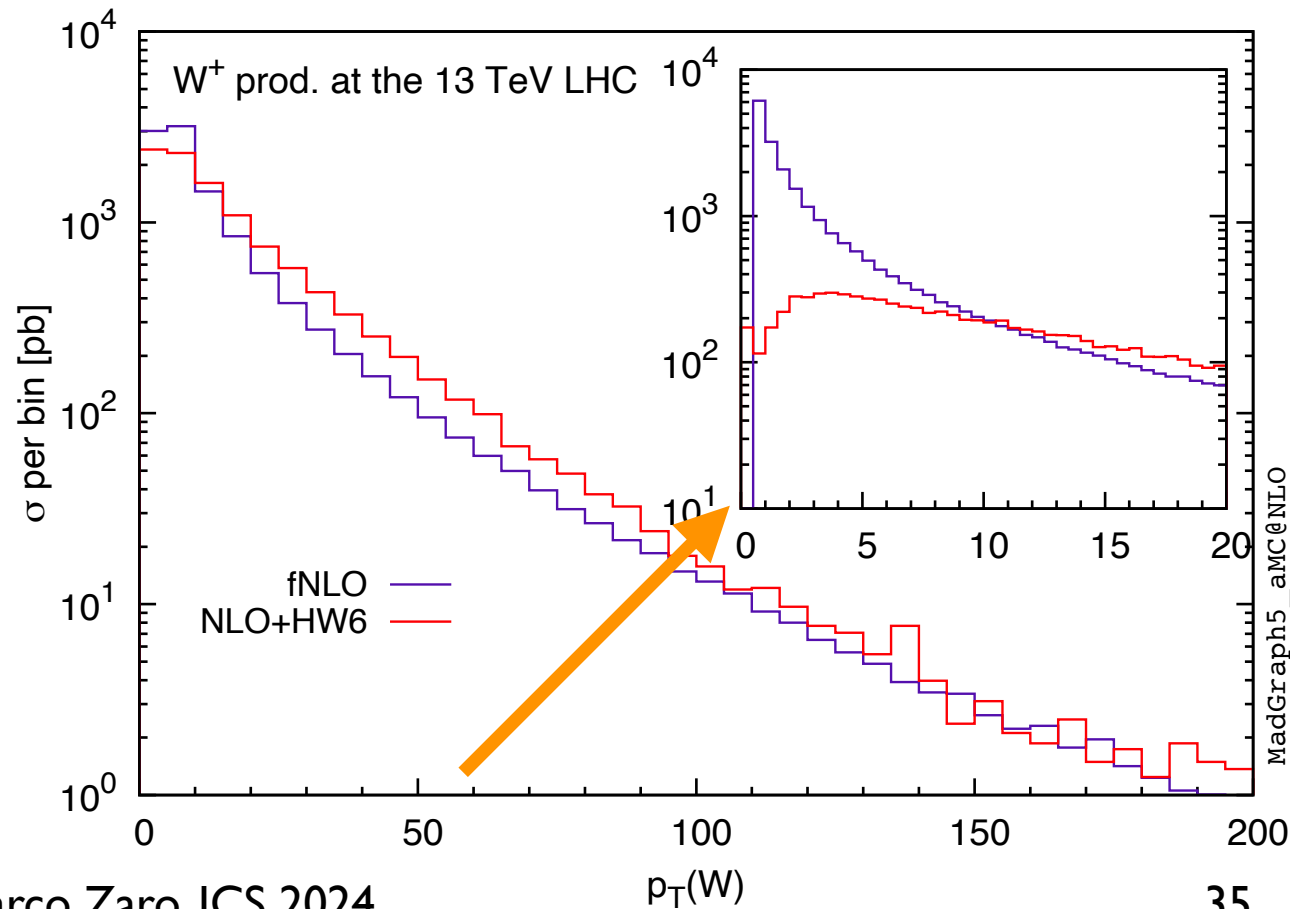
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- In practice, two set of momenta are generated during the MC integration
  - One (or more)  $n$ -body set(s), for Born, virtuals and counterterms
  - One  $n+1$ -body set, for the real emission
- The various terms are computed. Cuts are applied on the corresponding momenta and histograms are filled with the weight and kinematics of each term



# Instabilities at fixed order

- Besides the mis-binning problem, the kinematics mismatch can lead to odd behaviours of certain observables, in particular when some constraint coming from the  $n$ -body kinematics is relaxed in the  $n+1$ -body one





# Subtracting IR divergences: Summary

- Virtual and real matrix element are not finite, but their sum is. Subtraction methods can be used to extract divergences for real-emission matrix elements and cancel explicitly the poles from the virtuals
- Event and counterevents have different kinematics. Unweighting is not possible, we need to fill plots on-the-fly with weighted events
- For plots, only IR-safe observable with finite resolution must be used!



# Intermezzo: Is it all at NLO?

- Suppose we have a code for  $pp \rightarrow t\bar{t}$  @NLO. Are all the following (IR-safe) variables described at NLO?
  - top  $p_T$
  - $t\bar{t}$  pair  $p_T$
  - $t\bar{t}$  pair invariant mass
  - jet (extra parton)  $p_T$
  - $t\bar{t}$  azimuthal distance



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YES



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  - jet (extra parton)  $p_T$
  - $t\bar{t}$  azimuthal distance

YES  
NO



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  - $t\bar{t}$  pair invariant mass
  - jet (extra parton)  $p_T$
  - $t\bar{t}$  azimuthal distance

YES  
NO  
YES



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# Intermezzo: Is it all at NLO?

- Suppose we have a code for  $pp \rightarrow t\bar{t}$  @NLO. Are all the following (IR-safe) variables described at NLO?
  - top  $p_T$  YES
  - $t\bar{t}$  pair  $p_T$  NO
  - $t\bar{t}$  pair invariant mass YES
  - jet (extra parton)  $p_T$  NO
  - $t\bar{t}$  azimuthal distance



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• Suppose we have a code for  $pp \rightarrow t\bar{t}$  @NLO. Are all the following (IR-safe) variables described at NLO?

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- $t\bar{t}$  pair  $p_T$
- $t\bar{t}$  pair invariant mass
- jet (extra parton)  $p_T$
- $t\bar{t}$  azimuthal distance

YES

NO

YES

NO

NO

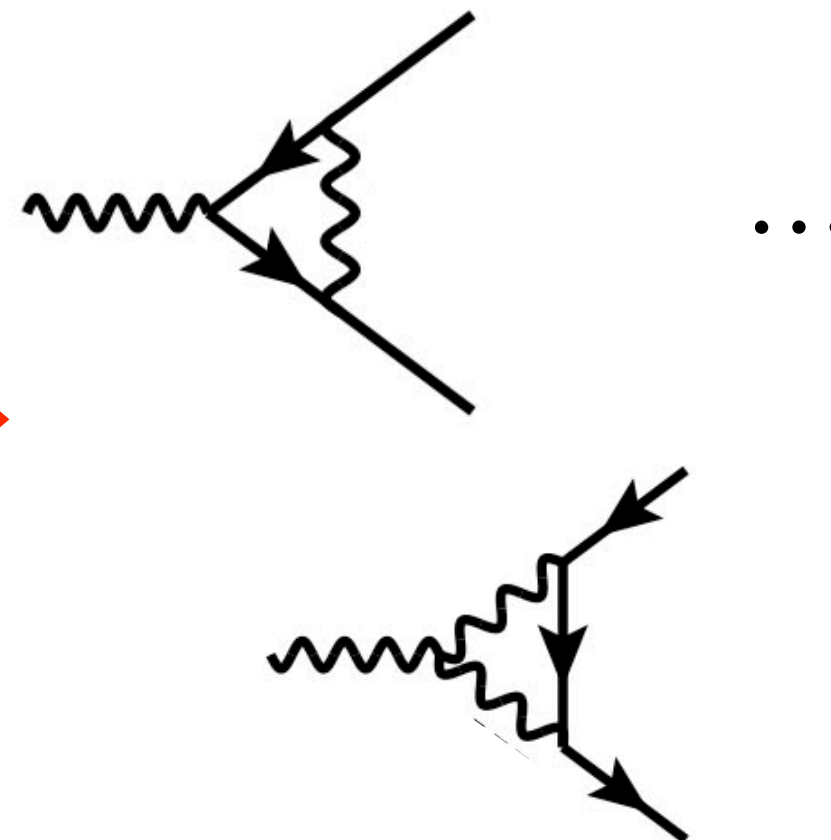
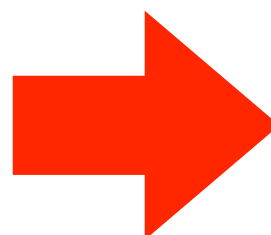
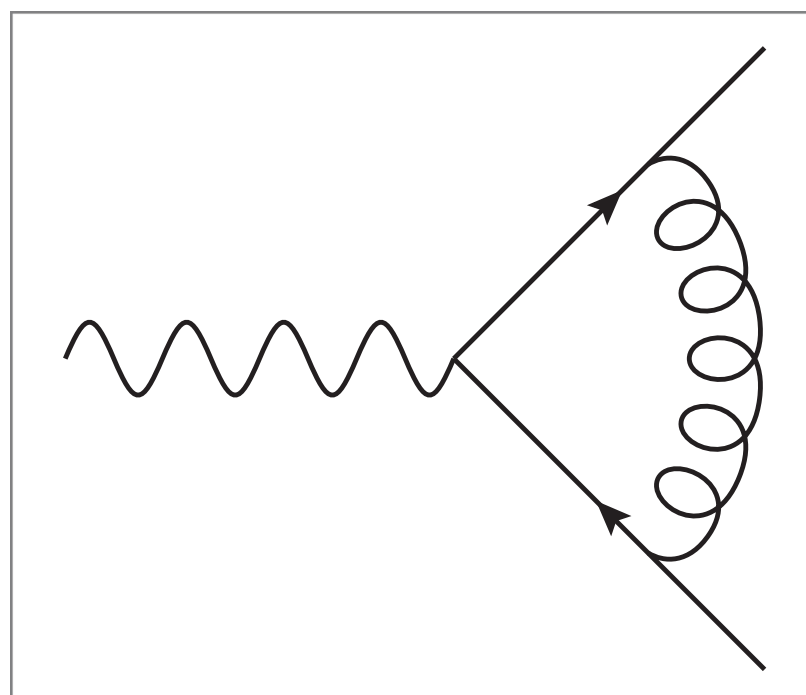


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# From QCD to EW corrections

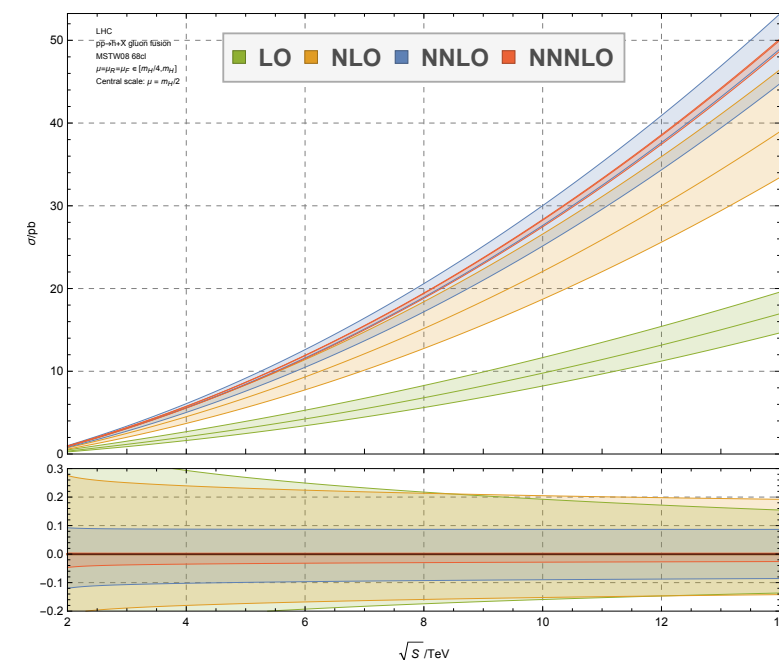
*a brief overview*



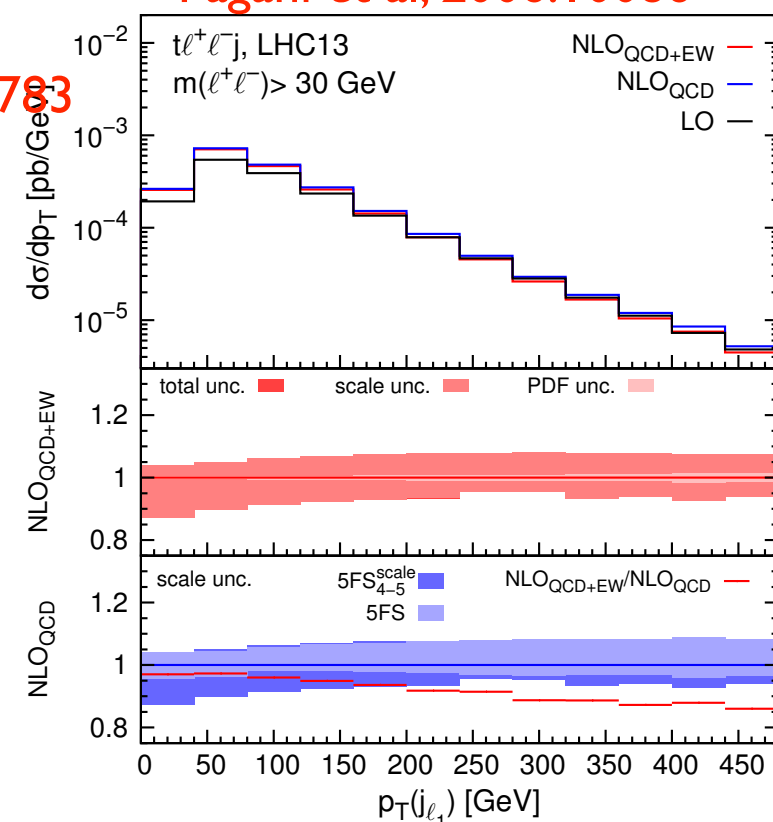
# Why bothering?

- QCD corrections generally improve precision of computations (shrink theoretical errors)
- EW corrections necessary to improve accuracy of predictions, specially in the tails of distributions (Sudakov enhancement)
- EW corrections are crucial at lepton colliders
- EW and complete-NLO corrections automated! [Sherpa+Openloops: 1412.5157](#); [Sherpa+Recola: 1704.05783](#); [MG5\\_aMC: 1804.10017](#)
- In some cases, EW corrections do not behave as expected: can give effects as large as QCD!

Anastasiou et al, 1503.06056



Pagani et al, 2006.10086

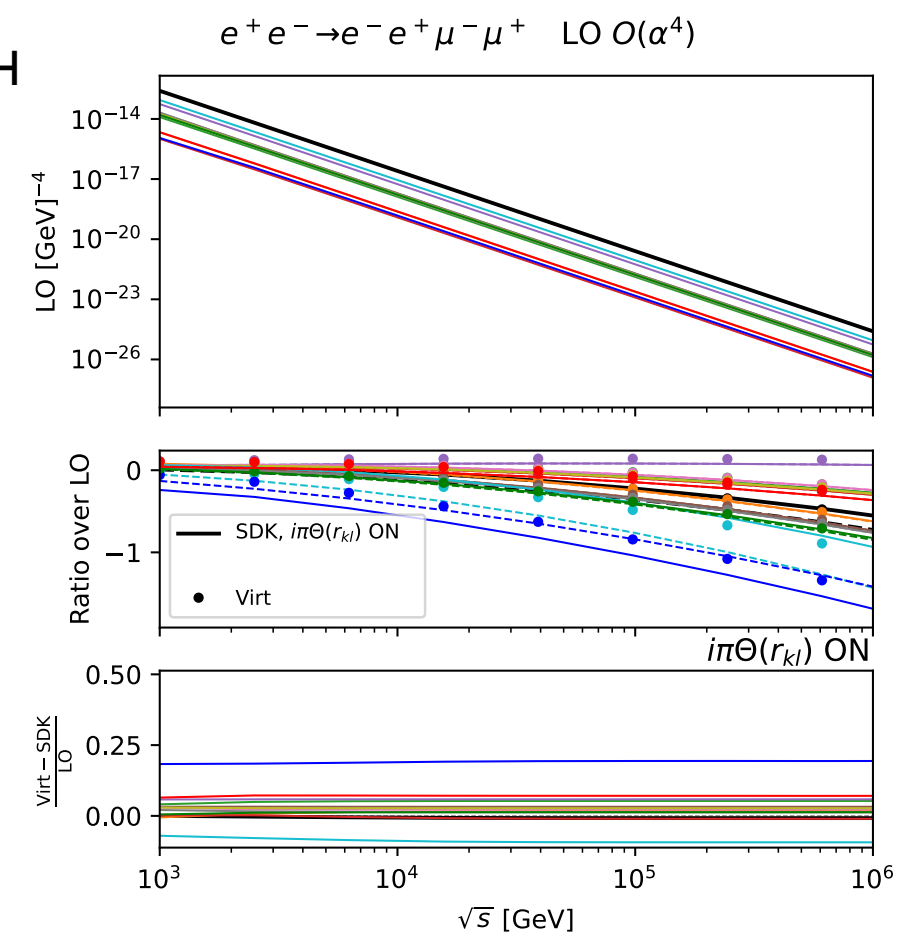
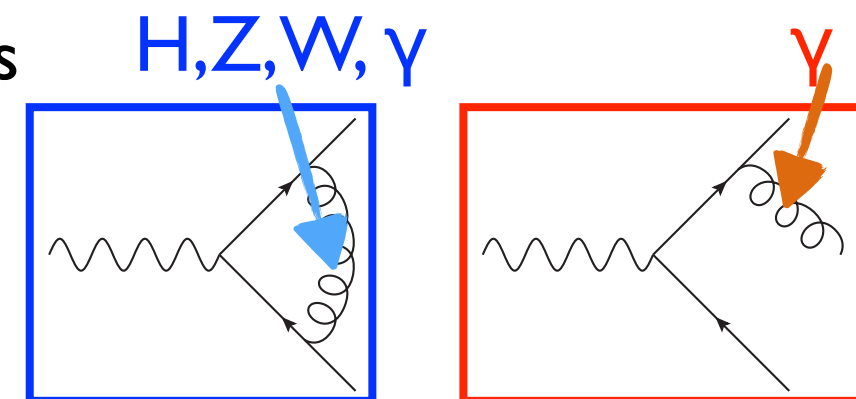


# Sudakov enhancement

Denner, Pozzorini, hep-ph/0010201 & hep-ph/0104127

Pagani, MZ, arXiv:2110.03714

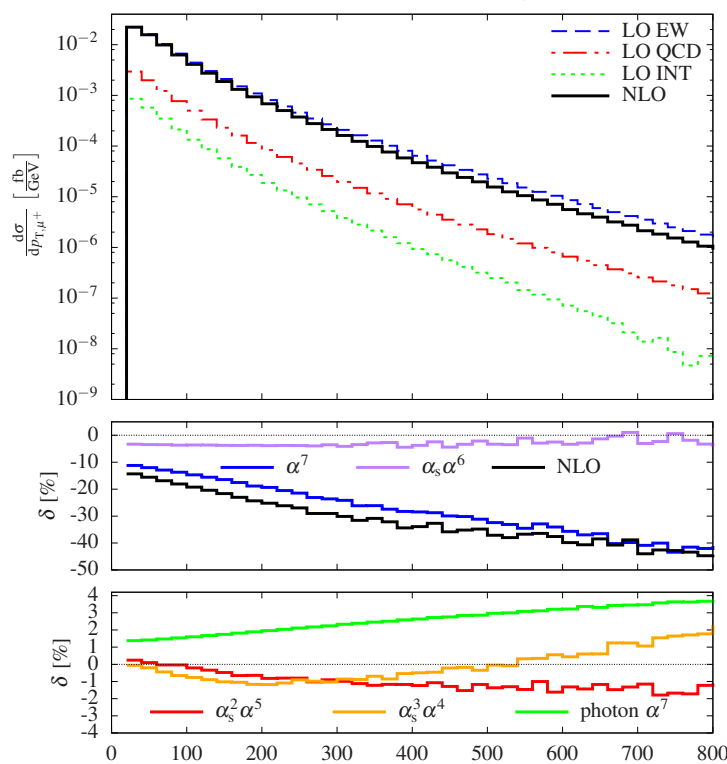
- EW bosons are massive: a real W/Z/Higgs emission is detectable (at least in principle)
- Radiation of W/Z/Higgs bosons is in general not included in EW corrections, which remain finite
- When the process scale  $Q$  is large,  $Q \gg M \sim m_W, m_Z, m_H$  the would-be IR divergence associated to the heavy boson shows up with double and single  $\log(Q/M)$
- In the regime where all invariants are  $\gg M$ , these logs are universal, and exponentiate at all orders (resummation possible)
- Sudakov approximation is excellent at high-energy (only a constant part is missing)



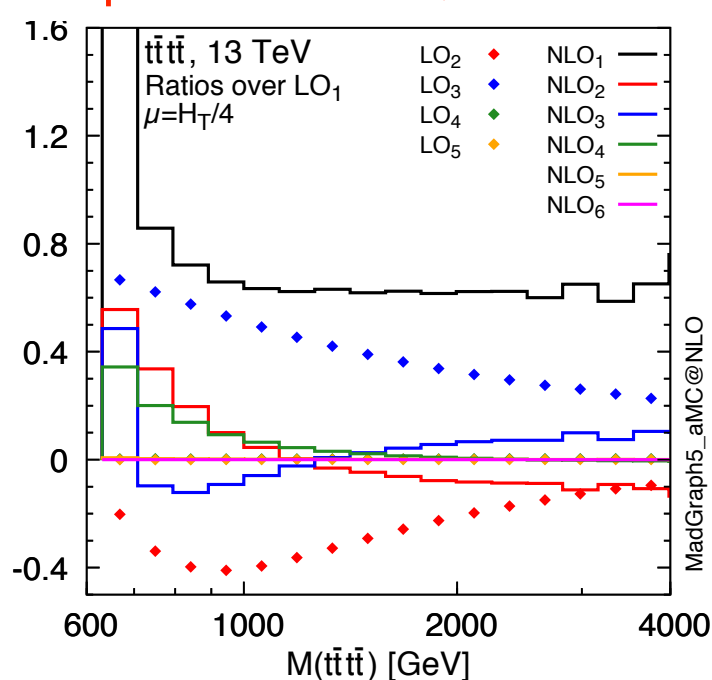
# Large EW corrections: not only Sudakov logs

- Despite the naive estimate  $\alpha \sim \alpha_s^2$ , there are cases when EW corrections comparable to NLO QCD or larger. It happens when:
  - Large scales are probed (VBS) feature of all VBS channels, see also Denner et al, 1904.00882, 2009.00411
  - Power counting is altered (4 top:  $\gamma_t$  vs  $\alpha$ )
  - New production mechanisms, different than those at the “dominant” LO, enter (ttW, bbH)

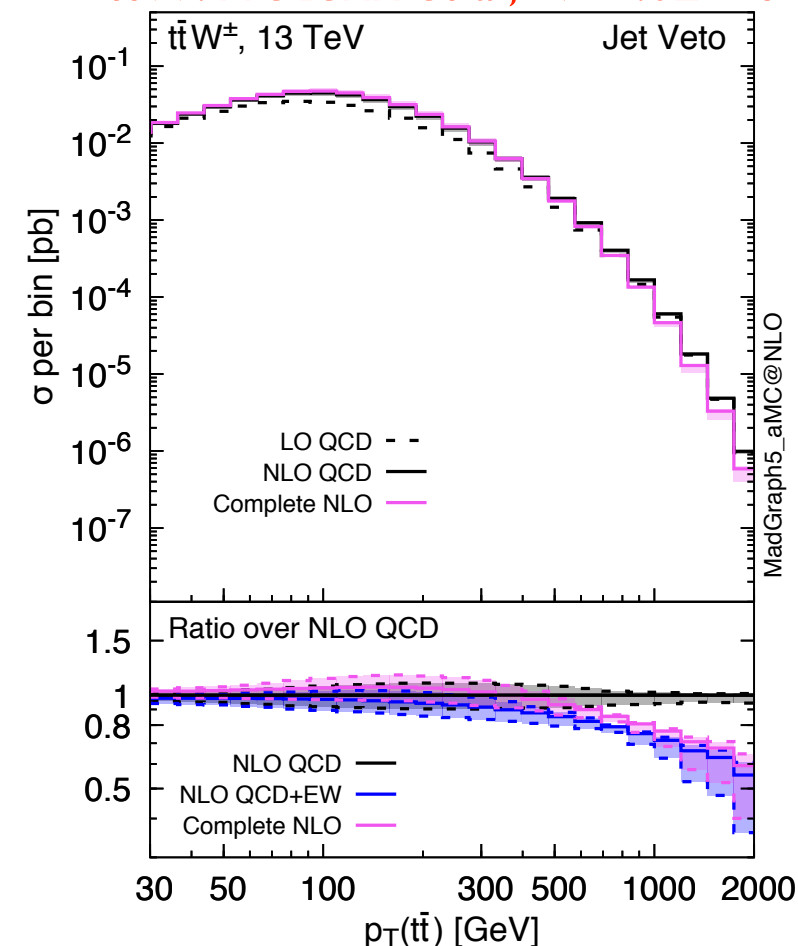
VBS: Biedermann et al, 1708.00268



4 top: Frederix et al, 1711.02116

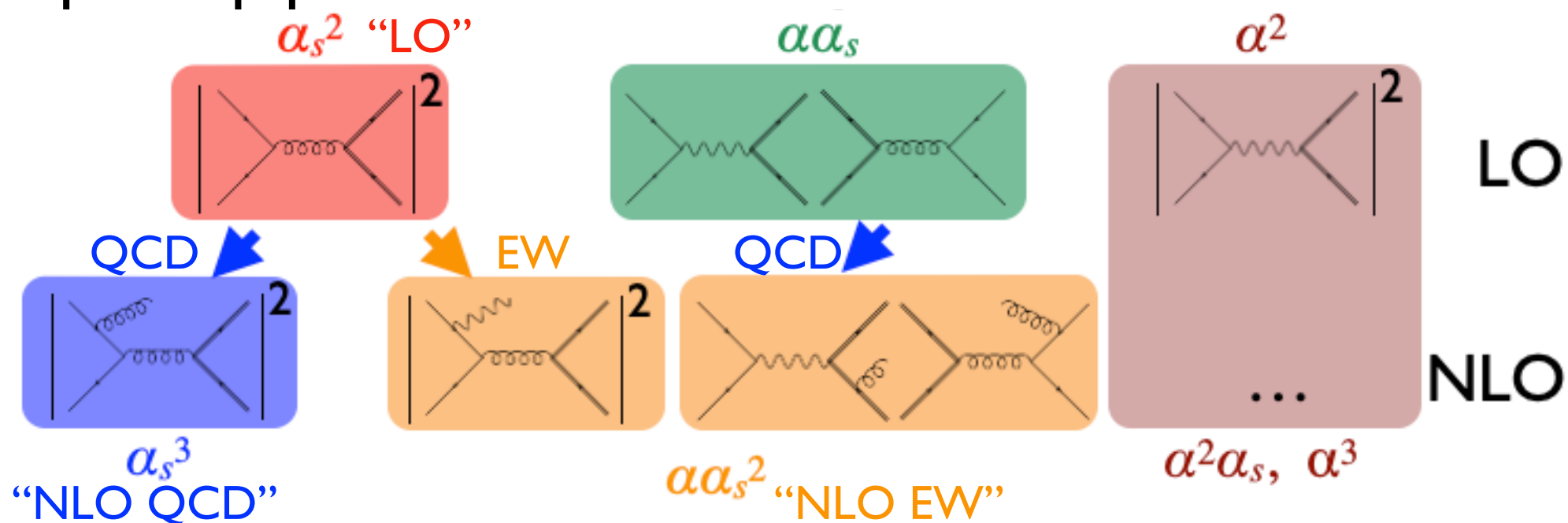


ttW: Frederix et al, 1711.02116



# Anatomy of EW corrections: EW corrections vs EW effects

- A general process has more contributions at LO, NLO, ...
- Example: top pair



- The **LO** is often identified with the contribution with most  $\alpha_s$
- At NLO the first two contributions are identified with the **NLO QCD** and **NLO EW** corrections
- This structure induces mixed QCD-EW effects at NLO:  

$$\text{NLO}_i = \text{LO}_{i-1} \otimes \text{EW} + \text{LO}_i \otimes \text{QCD}$$



# Multi-coupling expansion

Single coupling

$$\hat{\sigma} = \alpha_s^b \sigma_0 + \alpha_s^{b+1} \sigma_1 + \alpha_s^{b+2} \sigma_2 + \alpha_s^{b+3} \sigma_3 + \dots$$

LO      NLO      NNLO      NNNLO

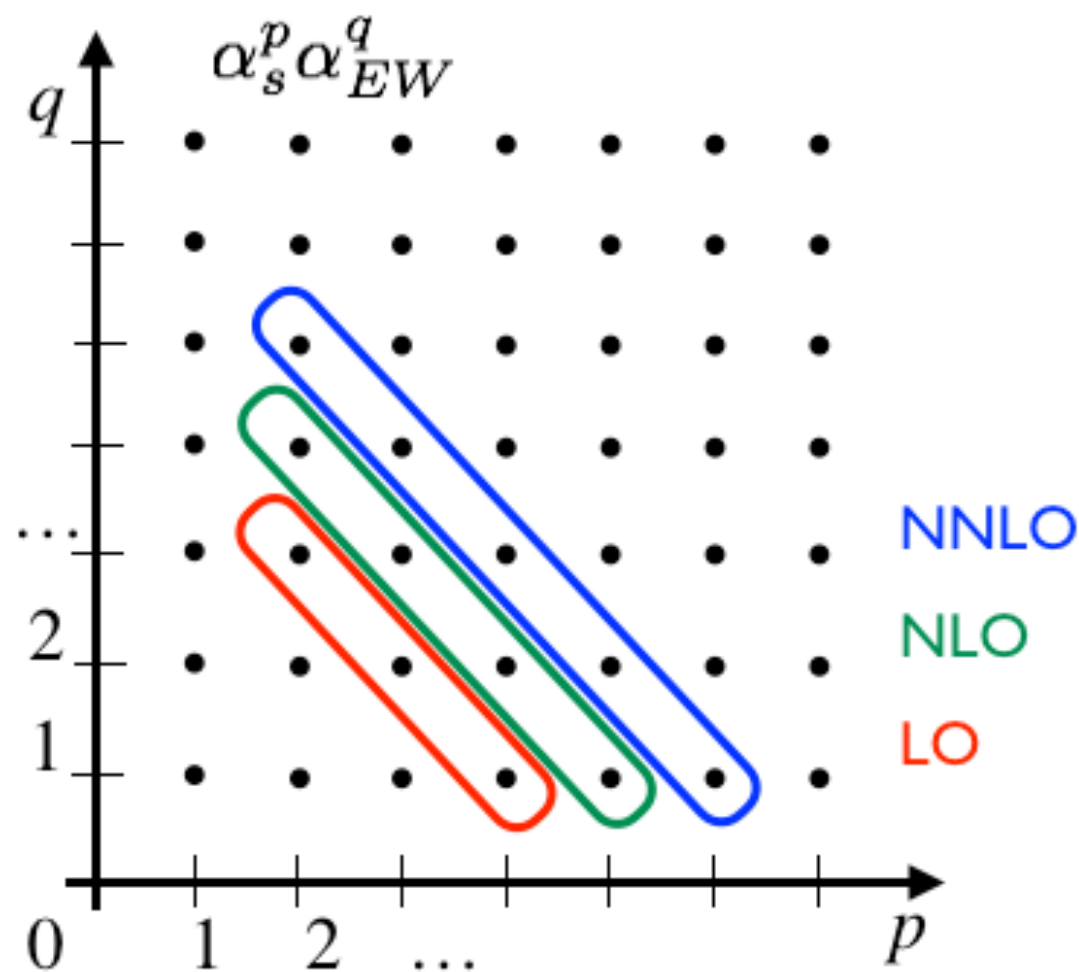
Multi-coupling

# Multi-coupling expansion

Single coupling  $\hat{\sigma} = \alpha_s^b \sigma_0 + \alpha_s^{b+1} \sigma_1 + \alpha_s^{b+2} \sigma_2 + \alpha_s^{b+3} \sigma_3 + \dots$

LO      NLO      NNLO      NNNLO

Multi-coupling







# Steps towards the automation of EW corrections

- Apart for the (much) more complex book-keeping, automation of NLO EW corrections largely builds on techniques for NLO QCD (modulo bookkeeping)
- IR subtraction: techniques established for QCD corrections can be extended to EW ones
- Replace color factors with charges ( $C_F \rightarrow q_i^2$ ,  $C_A \rightarrow 0$ ,  $T_F \rightarrow N_{C,i} q_i^2$ ) Replace color-linked Borns with charge-links
- Loop amplitudes: one-loop techniques can be exploited for EW loops.
- UV/R2 counterterms for the EW interactions are needed
- Higher ranks appear, integrand-reduction may lead to unstable results  
Switch to other techniques (Tensor-integral reduction, Laurent-series expansion,...)
- Use scalar-integral libraries that support complex masses





# EW renormalisation schemes in a nutshell

The renormalisation of  $\alpha$  can be performed in different schemes:

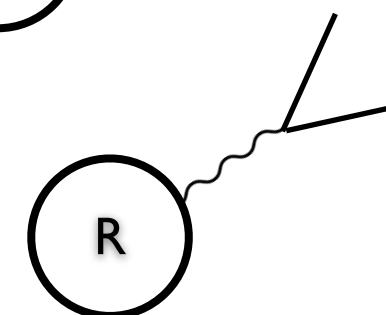
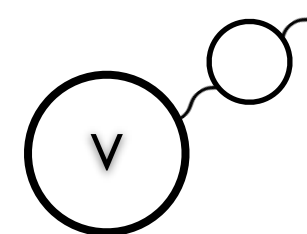
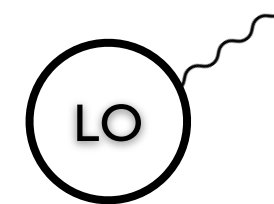
- $\alpha(0)$ :  $\alpha$  is measured in the Thompson scattering, in the zero-momentum limit. Terms  $\sim \log(Q/m_f)$  appear in the cross section, except for external photons. Fermion masses must be retained.
- $\alpha(M_Z)$ :  $\alpha$  is measured at the Z peak (e.g. at LEP). It removes the dependence on the fermion masses, which can be set to zero.
- $G_\mu$  scheme: the Fermi constant is measured from the muon lifetime, then  $\alpha$  is extracted. W.r.t. the  $\alpha(M_Z)$  scheme, also contributions of weak origin ( $\Delta\rho$ ) are resummed

The  $G_\mu$  scheme is generally preferred for processes without final-state photons at the LO.

# Processes with tagged photons

Pagani, Tsirikos, MZ arXiv:2106.02059

- The definition of a “photon” in the presence of EW corrections is not IR-safe (in a scheme with massless quarks/leptons)
- This is why democratic jets are usually employed
- In order to define photons as physical objects, a renormalisation scheme which takes into account fermion masses must be employed (only for the vertices related to tagged photons). Such a scheme exists:  $\alpha(0)$
- Renormalisation conditions define  $\alpha$  from the low-energy Thomson scattering. IR-poles differ from a high-energy scheme such as  $G_\mu$  or  $\alpha(m_Z)$
- The difference of IR poles accounts for the fact that real emissions with  $\gamma \rightarrow 2f$  splittings are not included
- Alternative: use fragmentation functions (more involved)





# NLO: Summary

- Precise predictions crucial for success of LHC programme
- They entail a lot of complexity: NLO is just the first bite!
- 10 years ago: NLO revolution. We have harvested many fruits
  - Automation: complexity hidden to the user!
  - NLO event generators ubiquitous in exp. analyses
  - Techniques proved successful also beyond QCD: automation of electroweak corrections (see backup slides for extra informations)



# Next?

- Beyond NLO: NNLO is the new Holy Graal:
  - Several subtraction techniques are being studied at NNLO. They all work on paper, need for numeric implementation and testing
  - No general algorithm to compute 2-loop amplitudes, but huge progress (first results for massless  $2 \rightarrow 3$  processes available)
  - In general, huge amount of complexity and of running time ( $\sim 1M$  CPU hours for  $2 \rightarrow 2$  with coloured FS)
- Is the NNLO revolution approaching?



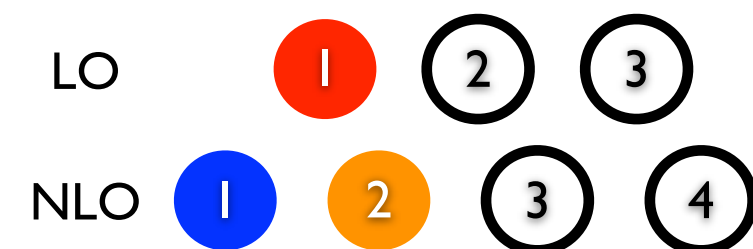
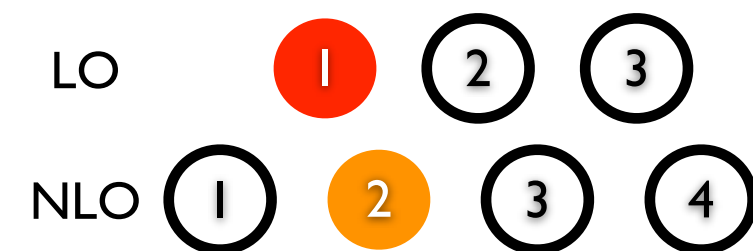
# Backup

# MG5\_aMC Syntax (I)

- The syntax to generate NLO EW corrections is very similar to the one for QCD:

- e.g.: `ttbar@NLO EW`: generate  $p p \rightarrow t t \sim$  [QED]
- Since no orders are specified, it will take the LO contribution with the largest power of  $\alpha_s^2$ ,  $\mathcal{O}(\alpha_s^2)$ , and generate NLO corrections with one extra power of  $\alpha$ ,  $\mathcal{O}(\alpha_s^2 \alpha)$

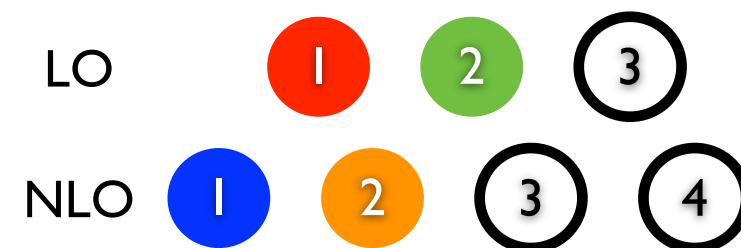
- If one wants to also generate NLO QCD corrections, the syntax is `generate p p > t t ~ [QED QCD]`  
In this case NLO contributions with both one extra power of  $\alpha$  and of  $\alpha_s$  will be generated





# MG5\_aMC Syntax (II)

- In the previous slide, the syntax would have been equivalent had we explicitly selected the dominant LO contribution.
- This could be done by adding  $QED^2=0$   $QCD^2=4$  to the generate command (note the squared-order constraints, applied at the amplitude level)
- Now, suppose you want to include also the first subleading LO term ( $LO_2$ ), together with NLO QCD and EW corrections.  
The syntax is: `generate p p > t t~ QED^2=2 QCD^2=4 [QCD]`.  
While counterintuitive, this is interpreted as in the previous slide:
- Generate LO contributions which satisfy the squared-order constraints ( $O(\alpha_s^2)$  and  $O(\alpha_s\alpha)$ )
- For the NLO corrections, add a power of  $\alpha_s$  on top of both. This will give ( $O(\alpha_s^3)$  and  $O(\alpha_s^2\alpha)$ )





# MG5\_aMC Syntax (III)

- Can I use diagram-order constraints?
- While this will give inconsistencies when NLO EW corrections are computed, it may be useful e.g. in EFT studies
- If the user asks for diagram constraints together with NLO corrections, the code will issue a clear warning, asking the user to acknowledge what he/she wants to do
- More info on <http://amcatnlo.cern.ch/co.htm>





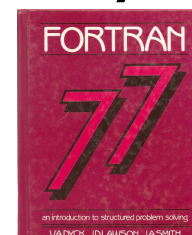
# Processes with tagged photons: how to

- In practice: a new model with both the HE renormalisation scheme ( $G_\mu$ ) and the  $\alpha(0)$  is available: `loop_qcd_qed_sm_Gmu-a0`
- Once loaded, tagged photons can be specified via the generate syntax:  
`generate t t~ !a! [QED]`
- Photons marked as tagged will not originate real emissions where  $\gamma \rightarrow 2f$  and the corresponding (local and integrated) FKS counterterms will not be included
- For each tagged photon, a term proportional to the difference between  $\alpha(0)$  and  $\alpha_{G_\mu}$  is added (it has IR poles)
- The final result is rescaled by  $(\alpha(0)/\alpha_{G_\mu})^{N_{\text{TagPhotons}}}$
- Result presented for top-pair and single-top production + photons  
[Pagani, Shao, Tsirikos, MZ 2106.02059](#)
- Available in v3.3.0

# Accessing the various coupling combinations

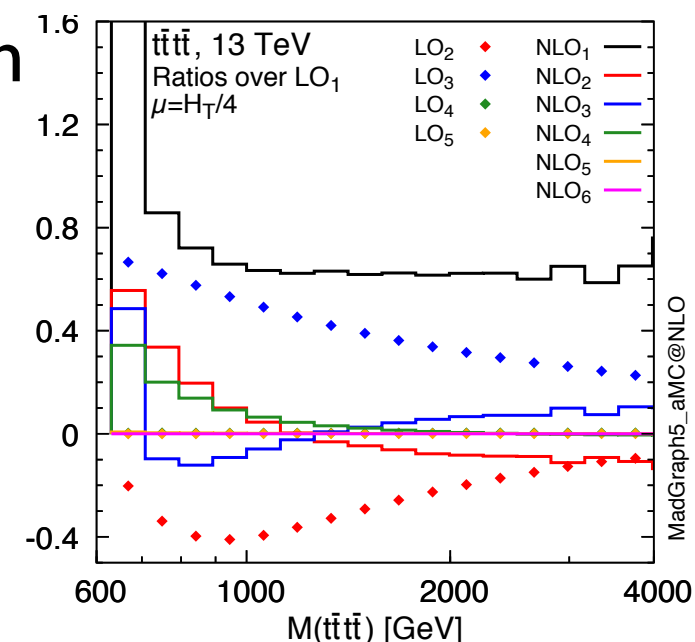
- The different coupling combinations to the cross section are evaluated in the same run
- Histograms can be booked for each of them in the analysis
- The coupling combination can be detected by using the `orders_tag_plot` variable

`integer orders_tag_plot`  
`common /corderstagplot/ orders_tag_plot`



- It is typically computed as  $100 \cdot \text{QED} + 1 \cdot \text{QCD}$  (may change if more coupling types are around)
- In any case, the specific values are printed inside the log file

```
INFO: orders_tag_plot is computed as:      + QCD *      1      + QED *      100
orders_tag_plot=      4  for QCD,QED, =      4 ,      0 ,
orders_tag_plot=     202  for QCD,QED, =      2 ,      2 ,
orders_tag_plot=     400  for QCD,QED, =      0 ,      4 ,
orders_tag_plot=      6  for QCD,QED, =      6 ,      0 ,
orders_tag_plot=     204  for QCD,QED, =      4 ,      2 ,
orders_tag_plot=     402  for QCD,QED, =      2 ,      4 ,
```





# Accessing the various coupling combinations in LHE events

- The same coupling structure can be accessed inside the LHE event file (when PS-matching is possible)
- Weights are stored in the same format as the scale/PDF variations

```

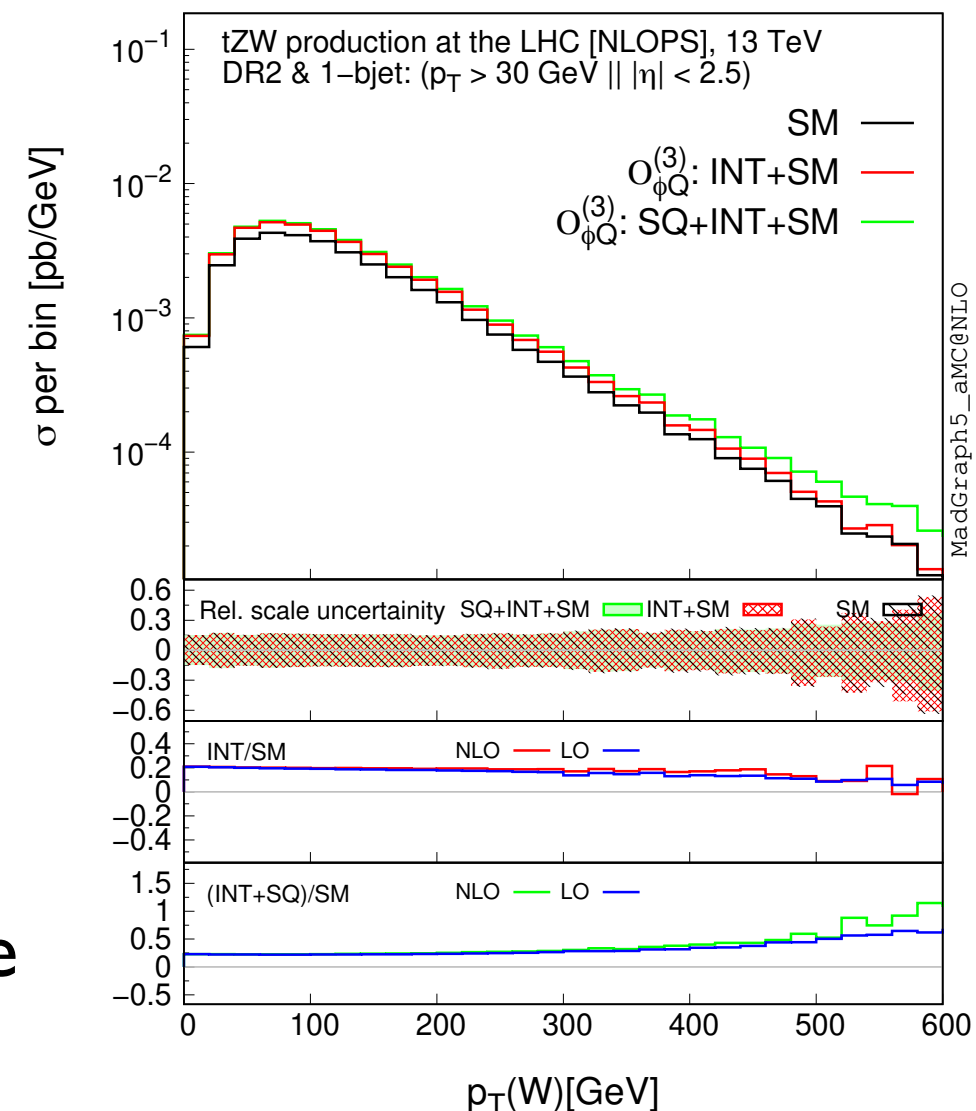
<initrwgt>
  <weightgroup name='scale_variation'          0  0' combine='envelope'>
    <weight id='1001'> tag=          0 dyn=  0 muR=0.10000E+01 muF=0.10000E+01 </weight>
    <weight id='1002'> tag=          0 dyn=  0 muR=0.20000E+01 muF=0.10000E+01 </weight>
    <weight id='1003'> tag=          0 dyn=  0 muR=0.50000E+00 muF=0.10000E+01 </weight>
    <weight id='1004'> tag=          0 dyn=  0 muR=0.10000E+01 muF=0.20000E+01 </weight>
    <weight id='1005'> tag=          0 dyn=  0 muR=0.20000E+01 muF=0.20000E+01 </weight>
    <weight id='1006'> tag=          0 dyn=  0 muR=0.50000E+00 muF=0.20000E+01 </weight>
    <weight id='1007'> tag=          0 dyn=  0 muR=0.10000E+01 muF=0.50000E+00 </weight>
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    <weight id='1018'> tag=          40200 dyn=  0 muR=0.50000E+00 muF=0.50000E+00 </weight>
  </weightgroup>
  <weightgroup name='scale_variation'          40202  0' combine='envelope'>
    <weight id='1019'> tag=          40202 dyn=  0 muR=0.10000E+01 muF=0.10000E+01 </weight>
  ...
</initrwgt>

<event>
  5          0 0.15776264E+00 0.21383348
          -5 -1  0  0  0  501 0.0
          21 -1  0  0  501 502 0.0
          -6  1  1  2  0  502 -0.4
          24  1  1  2  0  0 0.7
          23  1  1  2  0  0 -0.3
  #aMCatNLO 1  0  0  1  2 0.91081533E+
  0.00000000E+00
  <rwgt>
    <wgt id='1001'> 0.15776E+00 </wgt>
    <wgt id='1002'> 0.15496E+00 </wgt>
    <wgt id='1003'> 0.15846E+00 </wgt>
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    <wgt id='1011'> 0.12227E+00 </wgt>
    <wgt id='1012'> 0.14798E+00 </wgt>
    <wgt id='1013'> 0.13946E+00 </wgt>
    <wgt id='1014'> 0.12736E+00 </wgt>
    <wgt id='1015'> 0.15414E+00 </wgt>
  ...

```

# Accessing the various coupling combinations

- In either case, having all the couplings available from the same run makes them all statistically-correlated
- It is specially useful in the context of EFT studies, where different admixtures of new-physics can be morphed starting from the event weights
- Careful when matching to PS!  
If the statistical distribution of colour-flows is very different from one coupling combination to another (e.g. EFT vs SM), morphing could be dangerous!



El Faham, Maltoni, Mimasu, MZ  
arXiv:2111.03080